

This article was downloaded by:

On: 21 January 2011

Access details: *Access Details: Free Access*

Publisher *Taylor & Francis*

Informa Ltd Registered in England and Wales Registered Number: 1072954 Registered office: Mortimer House, 37-41 Mortimer Street, London W1T 3JH, UK



## International Reviews in Physical Chemistry

Publication details, including instructions for authors and subscription information:

<http://www.informaworld.com/smpp/title~content=t713724383>

### Electronic spectra of carbon chains and derivatives

Ramya Nagarajan<sup>a</sup>; John P. Maier<sup>a</sup>

<sup>a</sup> Department of Chemistry, University of Basel, Klingelbergstrasse 80, Basel, Switzerland

Online publication date: 16 July 2010

**To cite this Article** Nagarajan, Ramya and Maier, John P.(2010) 'Electronic spectra of carbon chains and derivatives', *International Reviews in Physical Chemistry*, 29: 3, 521 – 554

**To link to this Article:** DOI: 10.1080/0144235X.2010.490328

**URL:** <http://dx.doi.org/10.1080/0144235X.2010.490328>

PLEASE SCROLL DOWN FOR ARTICLE

Full terms and conditions of use: <http://www.informaworld.com/terms-and-conditions-of-access.pdf>

This article may be used for research, teaching and private study purposes. Any substantial or systematic reproduction, re-distribution, re-selling, loan or sub-licensing, systematic supply or distribution in any form to anyone is expressly forbidden.

The publisher does not give any warranty express or implied or make any representation that the contents will be complete or accurate or up to date. The accuracy of any instructions, formulae and drug doses should be independently verified with primary sources. The publisher shall not be liable for any loss, actions, claims, proceedings, demand or costs or damages whatsoever or howsoever caused arising directly or indirectly in connection with or arising out of the use of this material.

## Electronic spectra of carbon chains and derivatives

Ramya Nagarajan and John P. Maier\*

*Department of Chemistry, University of Basel,  
Klingelbergstrasse 80, Basel, Switzerland*

*(Received 6 April 2010; final version received 28 April 2010)*

Spectroscopic characterisation of carbon chains and ions has been motivated by a general interest in understanding the electronic structure and properties and also due to their prominence in astrochemistry, combustion processes as well as molecular electronics. Various experimental and theoretical approaches have been successfully implemented over the years towards realising this goal. In the current review, we have tabulated the experimentally available data on the excited electronic states of carbon chains and derivatives containing H, N, O, S and metals.

**Keywords:** carbon chains; heteroatom-doped carbon chains; ISM; diffuse interstellar bands

	Contents	PAGE
<b>1. Introduction</b>		522
<b>2. Bare carbon chains</b>		523
2.1. $C_n$ , $C_n^+$ and $C_n^-$		523
<b>3. Hydrogen terminated carbon chains</b>		528
3.1. $C_nH$ , $C_nH^+$ and $C_nH^-$		528
3.2. $HC_nH$ , $HC_nH^+$ and $HC_nH^-$		533
<b>4. Nitrogen and phosphorous terminated carbon chains</b>		535
4.1. $C_nN$ , $C_nN^+$ and $C_nN^-$		535
4.2. $HC_nN$ , $HC_nN^+$ and $HC_nN^-$		540
4.3. $NC_nN$ and $NC_nN^+$		541
4.4. $C_nP$ , $C_nP^+$ and $C_nP^-$		543
<b>5. Oxygen terminated carbon chains</b>		543
5.1. $C_nO$ , $C_nO^+$ and $C_nO^-$		543
5.2. $HC_nO$ , $HC_nO^+$ and $HC_nO^-$		544
5.3. $OC_nO$		544

---

\*Corresponding author. Email: j.p.maier@unibas.ch

<b>6. Sulphur terminated carbon chains</b>	545
6.1. $C_nS$ , $C_nS^+$ and $C_nS^-$	545
6.2. $HC_nS$ , $HC_nS^+$ and $HC_nS^-$	547
6.3. $SC_nS$ and $SC_nS^-$	548
<b>7. Metal terminated carbon chains</b>	548
<b>8. Outlook</b>	550
<b>Acknowledgement</b>	551
<b>References</b>	551

## 1. Introduction

The first signs of molecules in the universe appeared during the recombination era, which marked the transition from an ionised state to a predominantly neutral environment made up of H, He, D and Li. Radiative association reactions triggered the formation of simple molecules, such as  $H_2$ , LiH and HD as well as molecular ions like  $He_2^+$ ,  $HeH^+$  and  $HD^+$ . As the universe began to cool and expand, stars were formed from giant clouds of hydrogen, thereby setting the stage for the formation of heavier and complex molecules containing C, O, N, S, P, halides and metals. Until now, approximately 150 molecules have been detected in the interstellar medium (ISM). From the first molecule CH to the most recent detection of ethylformate ( $C_2H_5OCHO$ ) and *n*-propylcyanide ( $C_3H_7CN$ ) [1], it is clear that the interstellar chemistry is largely carbon dominated. These molecules are known to be confined to distinct environments like the circumstellar shells of carbon or oxygen rich stars, planetary nebula, diffuse and dense clouds [2,3]. Different formation mechanisms ranging from gas phase ion-molecule reactions and neutral–neutral collisions to adsorption–desorption processes on dust grains have been proposed [4,5]. The molecular composition of space and the reaction pathways leading to their formation remains an active area of research.

Molecules possess characteristic absorption/emission features generated by their electronic, vibrational and rotational motion; hence their detection in space and terrestrial environment relies on spectroscopy. Astronomical observations have been made possible by radio astronomy, infrared (IR) and optical telescopes. Over the years, laboratory experiments simulating the temperature, pressure and density conditions in space as well as computational tools have complemented the above observations.

Majority of the molecules detected in the interstellar and circumstellar medium are built up on a carbon–carbon framework. The ability of carbon to hybridise gives rise to an assortment of molecules ranging from the simplest  $C_2$ , to the longest conjugated system,  $HC_{11}N$  and the larger and more stable polycyclic aromatic hydrocarbons. The list includes cations and anions as well. In this review, we focus on the experimentally available data on the electronic spectroscopy of carbon chains and their derivatives containing H, N, O, S and metals. An understanding of their electronic spectra allows one to make direct comparisons with astrophysical observations. Relevant in this context is the role of carbon chains which have been postulated as carriers of diffuse interstellar bands (DIBs): a group of nearly 200 absorption features in the 400–900 nm region with half widths ranging from

1 to 200  $\text{cm}^{-1}$  [6,7]. Over the years, significant effort has been put forth in this direction by recording electronic spectra of mass selected carbon species in low-temperature matrices in conjunction with gas phase measurements [8,9]. Trends within a homologous series of carbon chains particularly with respect to the absorption wavelengths and oscillator strengths have emerged which have been instrumental in narrowing down the size and type molecule to be pursued as a potential DIB carrier. Until recently, the role of anions in the ISM was generally disregarded because efficient mechanisms for their formation were not known. Electron attachment to a polar neutral core via the dipole bound states (DBS) of anions (which are electronic states close to the detachment threshold) is now considered as a pathway to anion production in the ISM. Hydrogen terminated carbon chains,  $\text{C}_n\text{H}$  are known to have a high dipole moment and electron affinity; the corresponding  $\text{C}_4\text{H}^-$  [10],  $\text{C}_6\text{H}^-$  [11] and  $\text{C}_8\text{H}^-$  [12] were the first anions to be recently detected in space.

In the terrestrial environment, carbon chains play an important role in flame and soot chemistry [13]. Electron conducting properties which arise from conjugated  $\pi$  electron bonds in these chains (cumulenic or acetylenic structure) make them ideally suited for use as molecular wires in electronic systems [14,15]. From the perspective of fundamental research, it is interesting to study how the electronic structure, geometry and dynamics vary with the addition of different heteroatoms to bare carbon chains. An interesting example is the change in the ground state geometry from linear in the case of  $\text{HC}_n\text{N}$  systems to bent in  $\text{HC}_n\text{O}$  chains. Another feature is the alternation in molecular properties between even and odd members within a homologous series. For example, the ground state symmetry alternates between  $^1\Sigma^+$  and  $^3\Sigma^-$  in even-odd systems of the type  $\text{C}_n\text{X}$  ( $\text{X} = \text{C}, \text{O}, \text{S}$ ),  $\text{XC}_n\text{X}$  ( $\text{X} = \text{H}, \text{N}, \text{O}, \text{S}$ ) or  $\text{HC}_n\text{X}$  ( $\text{X} = \text{N}$ ). This behaviour is also manifested in the binding energies, ionisation potentials and electron affinities leading to the preferential formation of certain chains.

In this review, we tabulate the ground and excited electronic state symmetries and band origins of carbon chains and derivatives recorded in the gas phase and in neon or other matrix environments. It must be emphasised that many of these molecules have also been investigated in detail via microwave [16] and IR spectroscopies [17] and have been the subject of numerous theoretical studies [18].

## 2. Bare carbon chains

### 2.1. $\text{C}_n$ , $\text{C}_n^+$ and $\text{C}_n^-$

The structure and spectroscopic properties of bare carbon species and ions have been studied over the past several years. Calculations predict that linear cumulenic chains with  $^1\Sigma_g^+$  ground electronic states are the low energy forms for the odd members while for the even-membered species  $\text{C}_4$ ,  $\text{C}_6$  and  $\text{C}_8$ , the linear structure ( $^3\Sigma_g^-$ ) is nearly isoenergetic with the monocyclic ring isomer ( $^1\text{A}_g$ ) [19,20]. Under experimental conditions, it is the linear form which appears to be entropically favoured and is readily formed in the gas phase and cryogenic matrices. Rotationally resolved infrared spectra are available for linear  $\text{C}_n$  up to  $n = 13$  [21,22] in addition to the electronic spectra up to  $n = 21$  (Table 1). Cyclic isomers of  $\text{C}_6$ ,  $\text{C}_8$ , ...,  $\text{C}_{14}$  have been trapped in Ar/Ne matrices [23–26] while the ring forms of  $\text{C}_{14}$ ,  $\text{C}_{18}$  and  $\text{C}_{22}$  have been observed in the gas phase via a resonance enhanced multiphoton ionisation (REMPI) technique [27,28].

Table 1. Bare carbon chains:  $C_n$  species.

Species/electronic ground state	$n$	Origin ( $\text{cm}^{-1}$ )	Phase	Reference
$C_n$	2	$A^1\Pi_u$ 8268	Gas	J. Mol. Spectrosc. <b>131</b> , 261 (1988)
$n$ even X $^3\Sigma_g^-$		$D^1\Sigma_u^+$ 43,227		
$(n=2, X^1\Sigma_g^+)$	3	$A^1\Pi_u$ 24,685	Gas	J. Chem. Phys. <b>122</b> , 244308 (2005)
$n$ odd X $^1\Sigma_g^+$		$B^1\Delta_u$ 37,594–32,258	Ne matrix	J. Chem. Phys. <b>113</b> , 10999 (2000)
		$D^1\Delta_u$ 42,264		J. Chem. Phys. <b>125</b> , 244307 (2006)
		$^1\Sigma_u^+$ 62,500		J. Phys. Chem. A <b>106</b> , 5779 (2002)
	4	$^3\Sigma_u^-$ 26,323	Gas	J. Chem. Phys. <b>112</b> , 9777 (2000)
	5	$^1\Pi_u$ 19,599	Gas	J. Chem. Phys. <b>125</b> , 094308 (2006)
		$^1\Pi_u$ 22,470		
		$^1\Sigma_u^+$ 42,030		
	6	$1^3\Sigma_u^-$ 19,558	Ne matrix	J. Chem. Phys. <b>103</b> , 54 (1995)
		$2^3\Sigma_u^-$ 42,105	Gas	J. Chem. Phys. <b>125</b> , 094308 (2006)
	7	$^1\Pi_u$ 18,440	Ne matrix	J. Chem. Phys. <b>104</b> , 4954 (1996)
		$^1\Sigma_u^+$ 39,556		
	8	$1^3\Sigma_u^-$ 15,630	Ne matrix	J. Chem. Phys. <b>103</b> , 54 (1995)
		$2^3\Sigma_u^-$ 36,075	Gas	J. Chem. Phys. <b>125</b> , 094308 (2006)
	9	$^1\Sigma_u^+$ 34,700	Gas	J. Chem. Phys. <b>125</b> , 094308 (2006)
	10	$^3\Sigma_u^-$ 13,596	Ne matrix	J. Chem. Phys. <b>103</b> , 54 (1995)
	11	$^1\Sigma_u^+$ 29,732	Ne matrix	J. Chem. Phys. <b>104</b> , 4954 (1996)
	13	$^1\Sigma_u^+$ 26,341	Ne matrix	J. Chem. Phys. <b>104</b> , 4954 (1996)
	15	$^1\Sigma_u^+$ 23,832	Ne matrix	J. Chem. Phys. <b>104</b> , 4954 (1996)
	17	$^1\Sigma_u^+$ 21,719	Ne matrix	Chem. Phys. Lett. <b>304</b> , 35 (1999)
	19	$^1\Sigma_u^+$ 19,867	Ne matrix	Chem. Phys. Lett. <b>304</b> , 35 (1999)
	21	$^1\Sigma_u^+$ 18,387	Ne matrix	Chem. Phys. Lett. <b>304</b> , 35 (1999)
$C_n^+$	2	$B^4\Sigma_u^+$ 19,730	Gas	J. Chem. Phys. <b>88</b> , 4614 (1988)
$n$ even X $^2\Pi_{g/u}$	6	$^2\Pi_g$ 15,484	Ne matrix	J. Chem. Phys. <b>120</b> , 7520 (2004)
$(n=2, X^4\Sigma_g^+)$	7	A 12,985	Ne matrix	J. Chem. Phys. <b>123</b> , 044305 (2005)
$n$ odd X $^2\Pi_{g/u}$		B 30,093		
		C 32,395		
	8	A 11,226	Ne matrix	J. Chem. Phys. <b>123</b> , 044305 (2005)
		B 16,624		
		C 16,729		
		D 19,268		
		E 32,453		
	9	A 26,918	Ne matrix	J. Chem. Phys. <b>123</b> , 044305 (2005)
$C_n^-$	2	$A^2\Pi_u$ 3928	Gas	J. Chem. Phys. <b>89</b> , 129 (1988)
$n$ even X $^2\Pi_{g/u}$		$B^2\Sigma_u^+$ 18,483		Chem. Phys. Lett. <b>244</b> , 400 (1995)
$(n=2, X^2\Sigma_g^+)$	3	$A^2\Delta_u$ 4743	Gas	J. Chem. Phys. <b>112</b> , 3747 (2000)
		$B^2\Sigma_u^-$ 25,948		
		$C^2\Sigma_u^+$ 28,507		
$n$ odd X $^2\Pi_{g/u}$	4	$A^2\Sigma_g^+$ 8019	Ne matrix	J. Chem. Phys. <b>103</b> , 54 (1995)

(continued)

Table 1. Continued.

Species/electronic ground state	$n$	Origin ( $\text{cm}^{-1}$ )	Phase	Reference
$C_n^-$ $n$ even X $^2\Pi_{g/u}$ $n$ odd X $^2\Pi_{g/u}$	5	$B^2\Sigma_u^+$ 10,789	Gas	Chem. Phys. <b>228</b> , 293 (1998)
		$C^2\Pi_u$ 21,896		
		$2^2\Pi_u$ 26,069		
		$3^2\Pi_u$ 29,036		
		$C^2\Pi_u$ 25,539		
		$2^2\Pi_u$ 25,989		
		$3^2\Pi_u$ 28,717		
		$A^2\Pi_g$ 20,200	Ne matrix	J. Phys. Chem. A. <b>101</b> , 5292 (1997)
		$B^2\Pi_g$ 27,847	Gas	Chem. Phys. <b>228</b> , 293 (1998)
		$A^2\Pi_g$ 20,164		
$C_n^-$ $n$ even X $^2\Pi_{g/u}$ $n$ odd X $^2\Pi_{g/u}$	6	$A^2\Sigma_g^+$ 9352	Ne matrix	J. Chem. Phys. <b>103</b> , 54 (1995)
		$C^2\Pi_u$ 16,453	Gas	J. Chem. Phys. <b>105</b> , 4905 (1996)
		$2^2\Pi_u$ 20,064		
		$3^2\Pi_u$ 22,517		
		$C^2\Pi_u$ 16,476		
		$A^2\Pi_g$ 15,954	Ne matrix	J. Phys. Chem. A. <b>101</b> , 5292 (1997)
		$B^2\Pi_g$ 20,314	Gas	Chem. Phys. <b>228</b> , 293 (1998)
		$C^2\Pi_g$ 35,231		
		$A^2\Pi_g$ 15,943		
		$B^2\Pi_g$ 20,298		
$C_n^-$ $n$ even X $^2\Pi_{g/u}$ $n$ odd X $^2\Pi_{g/u}$	8	$^2\Sigma_u^+$ 1047	Ne matrix	J. Chem. Phys. <b>107</b> , 4468 (1997)
		$C^2\Pi_u$ 12,933	Gas	J. Chem. Phys. <b>105</b> , 4905 (1996)
		$2^2\Pi_u$ 16,295		
		$C^2\Pi_u$ 12,963		
		$2^2\Pi_u$ 16,305		
		$A^2\Pi_g$ 13,082	Ne matrix	J. Phys. Chem. A. <b>101</b> , 5292 (1997)
		$B^2\Pi_g$ 16,468	Gas	Chem. Phys. <b>228</b> , 293 (1998)
		$C^2\Pi_g$ 29,446		
		$D^2\Pi_g$ 34,536		
		$A^2\Pi_g$ 13,080		
$C_n^-$ $n$ even X $^2\Pi_{g/u}$ $n$ odd X $^2\Pi_{g/u}$	10	$B^2\Sigma_g^+$ 8964	Ne matrix	J. Chem. Phys. <b>107</b> , 4468 (1997)
		$C^2\Pi_u$ 10,338	Gas	J. Chem. Phys. <b>109</b> , 1329 (1998)
		$2^2\Pi_u$ 13,596		
		$C^2\Pi_u$ 10380		
		$A^2\Pi_g$ 11,042	Ne matrix	J. Phys. Chem. A. <b>101</b> , 5292 (1997)
		$B^2\Pi_g$ 13,906	Gas	J. Chem. Phys. <b>106</b> , 9992 (1997)
		$C^2\Pi_g$ 25,151		
		$D^2\Pi_g$ 29,884		

(continued)

Table 1. Continued.

Species/electronic ground state	<i>n</i>	Origin (cm <sup>-1</sup> )	Phase	Reference
		A <sup>2</sup> Π <sub>g</sub> 11,050		
	12	C <sup>2</sup> Π <sub>u</sub> 8006	Ne matrix	J. Chem. Phys. <b>107</b> , 4468 (1997)
		2 <sup>2</sup> Π <sub>u</sub> 11,636	Gas	J. Phys. Chem. A. <b>104</b> , 8622 (2000)
		C <sup>2</sup> Π <sub>u</sub> 8100		
	13	A <sup>2</sup> Π <sub>g</sub> 9502	Ne matrix	Chem. Phys. Lett. <b>304</b> , 35 (1999)
		A <sup>2</sup> Π <sub>g</sub> 9500	Gas	J. Chem. Phys. <b>106</b> , 9992 (1997)
		B <sup>2</sup> Π <sub>g</sub> 11,300		
	14	C <sup>2</sup> Π <sub>u</sub> 6849	Ne matrix	J. Chem. Phys. <b>107</b> , 4468 (1997)
		2 <sup>2</sup> Π <sub>u</sub> 10,202	Gas	J. Phys. Chem. A. <b>104</b> , 8622 (2000)
		C <sup>2</sup> Π <sub>u</sub> 6900		
		2 <sup>2</sup> Π <sub>u</sub> 10,200		
	15	A <sup>2</sup> Π <sub>g</sub> 8316	Ne matrix	Chem. Phys. Lett. <b>304</b> , 35 (1999)
		A <sup>2</sup> Π <sub>g</sub> 8300	Gas	J. Phys. Chem. A. <b>104</b> , 8622 (2000)
		B <sup>2</sup> Π <sub>g</sub> 10,100		
	16	C <sup>2</sup> Π <sub>u</sub> 5784	Ne matrix	J. Chem. Phys. <b>103</b> , 54 (1995)
		C <sup>2</sup> Π <sub>u</sub> 5800	Gas	J. Phys. Chem. A. <b>104</b> , 8622 (2000)
		2 <sup>2</sup> Π <sub>u</sub> 9100		
	17	A <sup>2</sup> Π <sub>g</sub> 7382	Ne matrix	Chem. Phys. Lett. <b>304</b> , 35 (1999)
		A <sup>2</sup> Π <sub>g</sub> 7300	Gas	J. Phys. Chem. A. <b>104</b> , 8622 (2000)
		B <sup>2</sup> Π <sub>g</sub> 9100		
	18	C <sup>2</sup> Π <sub>u</sub> 4833	Ne matrix	J. Chem. Phys. <b>103</b> , 54 (1995)
	19	A <sup>2</sup> Π <sub>g</sub> 6612	Ne matrix	Chem. Phys. Lett. <b>304</b> , 35 (1999)
		A <sup>2</sup> Π <sub>g</sub> 6600	Gas	J. Phys. Chem. A. <b>104</b> , 8622 (2000)
		B <sup>2</sup> Π <sub>g</sub> 8200		
	20	C <sup>2</sup> Π <sub>u</sub> 4099	Ne matrix	J. Chem. Phys. <b>103</b> , 54 (1995)
	21	A <sup>2</sup> Π <sub>g</sub> 5974	Ne matrix	Chem. Phys. Lett. <b>304</b> , 35 (1999)
		A <sup>2</sup> Π <sub>g</sub> 5900	Gas	J. Phys. Chem. A. <b>104</b> , 8622 (2000)
		B <sup>2</sup> Π <sub>g</sub> 7500		
	22	C <sup>2</sup> Π <sub>g</sub> 3449	Ne matrix	Chem. Phys. Lett. <b>304</b> , 35 (1999)

Structural information about carbon chain ions has been obtained by gas phase chromatography [29,30] and ion-molecule reactivity studies [31]. These suggest that species smaller than C<sub>10</sub><sup>-</sup> exclusively exist as linear chains while both isomers coexist for systems from C<sub>10</sub><sup>-</sup> to C<sub>30</sub><sup>-</sup>. In the case of cations, ion-mobility studies indicate that for C<sub>7</sub><sup>+</sup>-C<sub>10</sub><sup>+</sup> both the linear and ring structures coexist while only the latter is observed for larger species from C<sub>10</sub><sup>+</sup> to C<sub>20</sub><sup>+</sup>.

Linear carbon chains have been characterised using a variety of experimental techniques and information about the excited electronic states, vibrational frequencies, electron affinities and ionisation potentials have been determined. Experimental data on the excited electronic states of bare carbon chains are presented in Table 1.

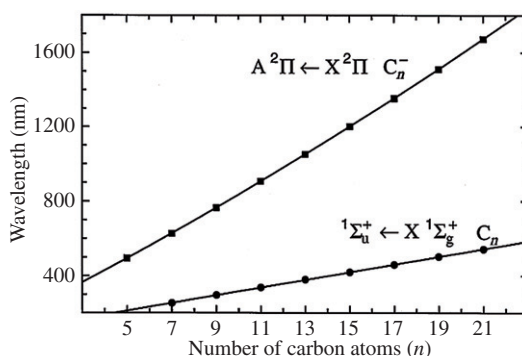


Figure 1. A plot of the origin band wavelength vs. the number of carbon atoms for the anionic and neutral odd-membered carbon chains.

The characteristic  $\pi \leftarrow \pi$  electronic excitation in these systems results in the  ${}^3\Sigma_u^- \leftarrow X^3\Sigma_g^-$  ( $n$  even) and  ${}^1\Sigma_u^+ \leftarrow X^1\Sigma_g^+$  ( $n$  odd) transitions. These were measured by depositing mass selected carbon anions in 5 K neon matrices followed by electron detachment with a broad band UV source. Optical spectra corresponding to the above transitions were observed only up to  $C_{10}$  in the even chains and farther up to  $C_{21}$  in the odd chains consistent with the other experimental observations which demonstrate a change in geometrical structure above  $C_{10}$ .

The matrix measurements have been the basis for gas phase studies of carbon chains.  $C_2$  and  $C_3$  are the best characterised of these. They have been detected in diffuse and translucent molecular clouds via their electronic transitions [32,33]. In addition,  $C_3$  and  $C_5$  have been detected in circumstellar shells of the carbon rich star, IRC+10216 by IR spectroscopy [34]. Consequently, a gas phase study of  $C_4$  was undertaken and a partly rotationally resolved spectrum of the origin band of the  ${}^3\Sigma_u^- \leftarrow X^3\Sigma_g^-$  transition was recorded using the cavity ring down (CRD) technique. Electronic spectra of  $C_5$ ,  $C_6$ ,  $C_8$  and  $C_9$  have been measured in the gas phase by REMPI spectroscopy. Figure 1 shows a plot of the origin band wavelength versus the number of carbon atoms for the  ${}^1\Sigma_u^+ \leftarrow X^1\Sigma_g^+$  and the  $A^2\Pi \leftarrow X^2\Pi$  band system of the neutral and anionic odd-membered carbon chains. The band origin shifts to the red with increasing chain length. In the case of  $C_5$ , three excited states were observed in which the spectral features were significantly broad and could not be rotationally resolved. This has been attributed to fast intramolecular processes occurring on a picosecond to femtosecond time scale. The intramolecular dynamics are depicted in Figure 2; upon electronic excitation, the molecule can either relax to higher vibrational levels of the ground state or get trapped in a nearby long-lived triplet state from where it can be subsequently ionised. The predissociative nature of the  $2^3\Sigma_u^-$  excited state of  $C_6$  and  $C_8$  and  ${}^1\Sigma_u^+$  of  $C_9$  was confirmed by a dissociative 'hole burning' approach in which a decrease in the ion current due to predissociation of the species of interest is monitored.

Carbon chain anions are open shell systems with  ${}^2\Pi_{g/u}$  ground states. The neutrals,  $C_n$  ( $n=2-11$ ) have high electron affinities, in the range of 2.8–4.4 eV [35], resulting in one or more bound excited electronic states below the photodetachment threshold.  $C_3^-$  with an electron affinity of 1.99 eV is the only anion, which does not possess a bound excited state.



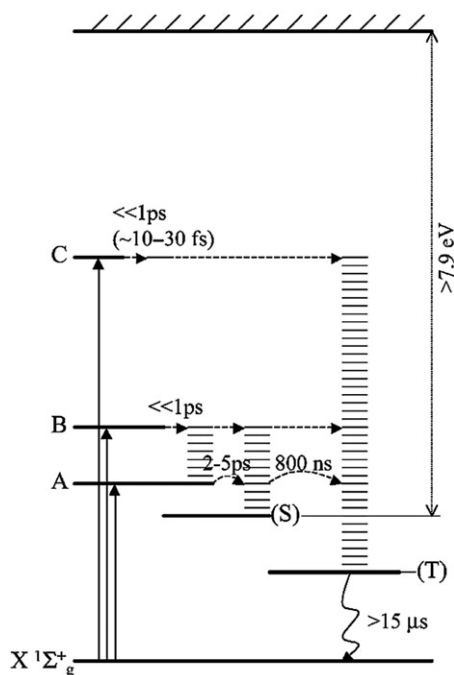


Figure 2. Inferred intramolecular dynamics in excited electronic states of C<sub>5</sub>.

UV-Vis absorption spectra of mass selected anions, C<sub>2n+1</sub><sup>-</sup> ( $n=2\text{--}10$ ) and C<sub>2n</sub><sup>-</sup> ( $n=2\text{--}11$ ) were first observed in rare gas matrices. Several electronic transitions were identified and assigned as <sup>2</sup>Π ← X <sup>2</sup>Π excitation. In addition, the <sup>2</sup>Σ<sup>+</sup> ← X <sup>2</sup>Π transitions have been observed for C<sub>2n</sub><sup>-</sup> ( $n=2\text{--}5$ ). Electronic spectra of anions in the gas phase have been measured using one- or two-colour resonant photodetachment technique. Excited states lying above the electron detachment threshold referred to as Feshbach states have been observed for C<sub>3</sub><sup>-</sup> and C<sub>5</sub><sup>-</sup>. Results from the matrix studies were used to assign several <sup>2</sup>Π ← X <sup>2</sup>Π transitions of the corresponding even and odd chains listed in Table 1.

In contrast to the neutral and anionic chains, experimental data on the electronic spectra of the cationic chains is limited. The B <sup>4</sup>Σ<sub>u</sub><sup>+</sup> ← X <sup>4</sup>Σ<sub>g</sub><sup>+</sup> band system of C<sub>2</sub><sup>+</sup> was observed in the gas phase and rotationally resolved by laser excitation spectroscopy. Recently, the <sup>2</sup>Π<sub>g</sub> ← X <sup>2</sup>Π<sub>g</sub> transition of C<sub>6</sub><sup>+</sup> and several dipole allowed transitions of linear C<sub>7</sub><sup>+</sup>, C<sub>8</sub><sup>+</sup> and C<sub>9</sub><sup>+</sup> have been observed in neon matrices along with their cyclic counterparts.

### 3. Hydrogen terminated carbon chains

#### 3.1. C<sub>n</sub>H, C<sub>n</sub>H<sup>+</sup> and C<sub>n</sub>H<sup>-</sup>

Hydrogen terminated carbon chain radicals, C<sub>n</sub>H ( $n=2\text{--}8$ ) have been detected in dark molecular clouds and envelopes of evolved stars [36]. The ground electronic states have been well characterised for C<sub>n</sub>H ( $n=2\text{--}11$ ) by microwave spectroscopy [37]. The dominant ground state electronic configuration for the even chains with  $n=2, 4$  is  $\pi^4\sigma^1$  giving <sup>2</sup>Σ<sup>+</sup>

symmetry; for  $n \geq 6$  it is  $\sigma^2\pi^3$  yielding a  ${}^2\Pi_i$  ground state. In the odd chains, the  $\pi$  highest occupied molecular orbital (HOMO) is singly occupied with  $\sigma^2\pi^1$  configuration resulting in a  ${}^2\Pi_r$  ground electronic state. These are known to exist in both cyclic and linear forms, with the latter becoming more stable with increasing  $n$ . For example, in  $C_3H$ ,  $C_5H$  and  $C_7H$ , the energy difference between the linear and cyclic isomers increases by 7.5, 15.5 and 21.3  $\text{kJ mol}^{-1}$ , respectively [38–40].

In the odd members, the lowest electronic states arise from  $\pi \leftarrow \sigma$  excitation giving  ${}^4\Sigma^-$ ,  ${}^2\Delta$ ,  ${}^2\Sigma^+$  and  ${}^2\Sigma^-$  excited states while  $\pi \leftarrow \pi$  electron promotion results in excited states with  ${}^2\Phi$ ,  ${}^4\Pi$  and  ${}^2\Pi(3)$  symmetries. The dipole allowed states include  ${}^2\Delta$ ,  ${}^2\Sigma^+$ ,  ${}^2\Sigma^-$  and  ${}^2\Pi(3)$ . Optical spectra have been recorded for  $C_nH$  ( $n=3, 5, 7$  and  $9$ ) by a REMPI technique. In the case of  $C_3H$ , a complex vibronic spectrum with a Franck-Condon profile showing a large change in geometry was observed and in conjunction with *ab initio* calculations the spectral features were assigned to linear–bent excitations. Three trans-bent excited states were observed with  $A^2A'$  and  $B^2A''$  correlating to  ${}^2\Delta$  in the linear configuration and a  $C^2A''$  state higher in energy. Unlike  $C_3H$ , the electronic spectra of the longer chains did not show a large geometry change and were assigned to linear–linear excitations. The band origins for the observed dipole allowed transitions are listed in Table 2, Panel A. The lowest excited states,  $A^2\Delta$  and  $B^2\Sigma^+$  for  $C_9H$  were not observed in the REMPI spectrum, which was recorded using 7.9 eV photons as the ionisation source; ionisation potential of  $C_9H$  is expected to be greater than 10.2 eV. However, transitions to these states are predicted to lie around 660 and 595 nm, respectively.

In the even numbered chains, the electronic spectra show signatures of vibronic coupling between the low-lying  ${}^2\Sigma^+$  and  ${}^2\Pi$  states. The  ${}^2\Pi$  state progressively becomes lower in energy as the chain length increases. The energy spacing between the two states has been deduced from the photoelectron spectra of the corresponding anions. In  $C_2H$  and  $C_4H$ , the  ${}^2\Pi$ – ${}^2\Sigma^+$  spacing is around 0.46 and 0.06 eV while the  ${}^2\Sigma^+$ – ${}^2\Pi$  separation for  $C_6H$  and  $C_8H$  is around 0.18 and 0.25 eV, respectively [41]. The second excited state,  $B^2\Pi$  has been observed in the gas phase for chains up to  $n=10$  and in neon matrices up to  $n=16$ .

The even chain anions are closed shell species with  ${}^1\Sigma^+$  ( $\pi^4$ ) ground electronic state, while the odd members exhibit  ${}^3\Sigma^-$  ( $\pi^2$ ) symmetry.  $\pi \leftarrow \pi$  excitation in both systems gives rise to  ${}^{1,3}\Sigma^+$ ,  ${}^{1,3}\Sigma^-$  and  ${}^{1,3}\Delta$  excited states. Electronic absorption spectra of the dipole allowed  ${}^1\Sigma^+ \leftarrow X^1\Sigma^+$  transition in the even series,  $C_nH^-$  ( $n=10, 12, \dots, 24$ ) and  ${}^3\Sigma^- \leftarrow X^3\Sigma^-$  in the odd series,  $C_nH^-$  ( $n=9–15$ ) have been observed; the spectral features assigned to two isomers. The first being the more stable linear acetylenic chain while the second a bent cumulene-like isomer with a quasilinear carbon chain framework and an off axis terminal hydrogen atom. For the second isomer, the bands were assigned to a  ${}^1A' \leftarrow X^1A'$  transition.

Excited electronic states close to the electron detachment threshold have been observed for the smaller anions,  $C_3H^-$ ,  $C_4H^-$ ,  $C_6H^-$  and  $C_8H^-$ . These are the DBS where the electron is weakly bound to the dipole field of the neutral [42]. A high electron affinity and dipole moment ( $>2D$ ) of the corresponding neutral are considered as prerequisites for their formation. Electronic transitions to the DBS of the above anions were observed by resonance enhanced multiphoton dissociation spectroscopy and assigned to  ${}^1\Pi \leftarrow X^1\Sigma^+$  of  $C_6H^-$  and  $C_8H^-$ , and  ${}^1A'({}^1\Pi) \leftarrow X^1\Sigma^+$  of  $C_4H^-$ . In  $C_6H^-$  and  $C_8H^-$ , the DBS arise from the  $X^2\Pi$  configuration of the neutrals which have dipole moments of 4.79 and 5.26 D [43], respectively. In the case of  $C_4H^-$ , the 0.8 D dipole moment of  $X^2\Sigma^+$  [44] is too low to

Table 2. H terminated carbon chains.

Species/electronic ground state	$n$	Origin (cm <sup>-1</sup> )	Phase	Reference
<i>Panel A: C<sub>n</sub>H species</i>				
C <sub>n</sub> H	2	A <sup>2</sup> Π 3750 B <sup>2</sup> Σ <sup>+</sup> 24,033	Gas	J. Chem. Phys. <b>108</b> , 10018 (1998)
$n$ even X <sup>2</sup> Σ <sup>+</sup> ( $n=2,4$ )				J. Mol. Spectrosc. <b>217</b> , 249 (2003)
X <sup>2</sup> Π <sub>i</sub> ( $n ≥ 6$ )	3	A <sup>2</sup> A' 19,187 B <sup>2</sup> A'' 20,538 C <sup>2</sup> A'' 23,171	Gas	J. Chem. Phys. <b>115</b> , 6913 (2001)
$n$ odd X <sup>2</sup> Π <sub>r</sub>	4	A <sup>2</sup> Π 468 B <sup>2</sup> Σ <sup>+</sup> 24,033	Gas	J. Chem. Phys. <b>108</b> , 10018 (1998) J. Chem. Phys. <b>108</b> , 3465 (1998)
	5	A <sup>2</sup> Δ 18,791 B <sup>2</sup> Σ <sup>-</sup> 20,608	Gas	J. Chem. Phys. <b>117</b> , 8362 (2002)
	6	A <sup>2</sup> Σ <sup>+</sup> 1460 B <sup>2</sup> Π 18,990	Gas	J. Chem. Phys. <b>108</b> , 10018 (1998) Faraday Diss. <b>109</b> , 109 (1998)
	7	A <sup>2</sup> Δ 16,819 B <sup>2</sup> Σ <sup>-</sup> 18,532 C <sup>2</sup> Σ <sup>+</sup> 20,627	Gas	J. Chem. Phys. <b>117</b> , 8362 (2002)
	8	A <sup>2</sup> Σ <sup>+</sup> 2041 B <sup>2</sup> Π 15,973	Gas	J. Chem. Phys. <b>108</b> , 10018 (1998) Chem. Phys. Lett. <b>382</b> , 245 (2003)
	9	C <sup>2</sup> Σ <sup>+</sup> 18,799 D <sup>2</sup> Π 20,205 F <sup>2</sup> Π 23,509	Gas	J. Chem. Phys. <b>117</b> , 8362 (2002)
	10	B <sup>2</sup> Π 14,003	Gas	J. Chem. Phys. <b>109</b> , 3819 (1998)
	12	B <sup>2</sup> Π 12,492	Ne matrix	J. Chem. Phys. <b>103</b> , 54 (1995)
	14	B <sup>2</sup> Π 11,554	Ne matrix	J. Chem. Phys. <b>103</b> , 54 (1995)
	16	B <sup>2</sup> Π 10,821	Ne matrix	J. Chem. Phys. <b>103</b> , 54 (1995)
C <sub>n</sub> H <sup>+</sup>	6	<sup>3</sup> Σ <sup>-</sup> 15,913	Ne matrix	J. Phys. Chem. A. <b>110</b> , 2885 (2006)
$n$ even X <sup>3</sup> Σ <sup>-</sup>	8	<sup>3</sup> Σ <sup>-</sup> 19,387	Ne matrix	J. Phys. Chem. A. <b>110</b> , 2885 (2006)
C <sub>n</sub> H <sup>-</sup>	3	d <sup>3</sup> A'' ← a <sup>3</sup> A''	Gas	Mol. Phys. <b>102</b> , 1881 (2004)
$n$ even X <sup>1</sup> Σ <sup>+</sup>		18,042		
$n$ odd X <sup>3</sup> Σ <sup>-</sup>	4	<sup>1</sup> A' ( <sup>1</sup> Π) 28,710 <sup>1</sup> A' ( <sup>1</sup> Π) 28,643	Gas	J. Chem. Phys. <b>108</b> , 10018 (1998) Mol. Phys. <b>101</b> , 583 (2003)
Acetylenic isomer (a)				
<sup>3</sup> Σ <sup>-</sup> ← X <sup>3</sup> Σ <sup>-</sup>	6	<sup>1</sup> Π 30,408	Gas	J. Chem. Phys. <b>116</b> , 6126 (2002)
<sup>1</sup> Σ <sup>+</sup> ← X <sup>1</sup> Σ <sup>+</sup>	8	<sup>1</sup> Σ <sup>+</sup> 35,174 <sup>1</sup> Π 31,497	Ne matrix Gas	J. Chem. Phys. <b>110</b> , 1492 (1999) J. Chem. Phys. <b>116</b> , 6126 (2002)
Cumulenic isomer (c)				
A' ← X <sup>1</sup> A'	9	<sup>3</sup> Σ <sup>-</sup> 27,630 (a) <sup>1</sup> A 29,014 (c)	Gas	J. Phys. Chem. A. <b>103</b> , 9712 (1999)
	10	<sup>1</sup> Σ <sup>+</sup> 30,479 (a)	Gas	J. Chem. Phys. <b>111</b> , 9280 (1999)

(continued)

Table 2. Continued.

Species/electronic ground state	<i>n</i>	Origin (cm <sup>-1</sup> )	Phase	Reference
		<sup>1</sup> A' 27,847 (c)	Ne matrix	
	11	<sup>3</sup> Σ <sup>-</sup> 24,288 (a)	Gas	J. Phys. Chem. A. <b>103</b> , 9712 (1999)
		<sup>1</sup> A' 26,742 (c)		
	12	<sup>1</sup> Σ <sup>+</sup> 26,802 (a)	Gas	J. Chem. Phys. <b>111</b> , 9280 (1999)
		<sup>1</sup> A' 25,654 (c)	Ne matrix	
	13	<sup>3</sup> Σ <sup>-</sup> 24,288 (a)	Gas	J. Phys. Chem. A. <b>103</b> , 9712 (1999)
		<sup>1</sup> A' 26,742 (c)		
	14	23,837 (a + c)	Ne matrix	J. Chem. Phys. <b>111</b> , 9280 (1999)
		23,763 (a + c)	Gas	
	15	<sup>3</sup> Σ <sup>-</sup> 19,434 (a)	Gas	J. Phys. Chem. A. <b>103</b> , 9712 (1999)
		<sup>1</sup> A' 23,284 (c)		
	16	<sup>1</sup> Σ <sup>+</sup> 21,226 (a)	Gas	J. Chem. Phys. <b>111</b> , 9280 (1999)
		<sup>1</sup> A' 21,852 (c)	Ne matrix	
	18	<sup>1</sup> Σ <sup>+</sup> 19,226 (a)	Gas	J. Chem. Phys. <b>111</b> , 9280 (1999)
		<sup>1</sup> A' 20,152 (c)	Ne matrix	
	20	<sup>1</sup> Σ <sup>+</sup> 17,592 (a)	Gas	J. Chem. Phys. <b>111</b> , 9280 (1999)
		<sup>1</sup> A' 18,641 (c)	Ne matrix	
	22	<sup>1</sup> A' 17,647 (c)	Ne matrix	J. Chem. Phys. <b>111</b> , 9280 (1999)
	24	<sup>1</sup> A' 16,731 (c)	Ne matrix	J. Chem. Phys. <b>111</b> , 9280 (1999)
<i>Panel B: HC<sub>n</sub>H species</i>				
HC <sub><i>n</i></sub> H	2	B <sup>1</sup> Σ <sub>u</sub> <sup>+</sup> 74,962	Gas	Helv. Chim. Acta. <b>57</b> , 1731 (1974)
<i>n</i> even X <sup>1</sup> Σ <sub>g</sub> <sup>+</sup>	4	B <sup>1</sup> Σ <sub>u</sub> <sup>+</sup> 60,864	Gas	Helv. Chim. Acta. <b>57</b> , 1731 (1974)
<i>n</i> odd X <sup>3</sup> Σ <sub>g</sub> <sup>-</sup>	6	B <sup>1</sup> Σ <sub>u</sub> <sup>+</sup> 54,615	Gas	Helv. Chim. Acta. <b>57</b> , 1731 (1974)
		A <sup>1</sup> Δ <sub>u</sub> 34,091		Phys. Chem. Chem. Phys. <b>5</b> , 4772 (2003)
	8	B <sup>1</sup> Σ <sub>u</sub> <sup>+</sup> 48,239	Gas	Helv. Chim. Acta. <b>57</b> , 1731 (1974)
		A <sup>1</sup> Δ <sub>u</sub> 29,454		Phys. Chem. Chem. Phys. <b>5</b> , 4772 (2003)
	10	B <sup>1</sup> Σ <sub>u</sub> <sup>+</sup> 43,784	Gas	J. Phys. Chem. A. <b>113</b> , 11099 (2009)
		A <sup>1</sup> Δ <sub>u</sub> 26,894		Phys. Chem. Chem. Phys. <b>5</b> , 4772 (2003)
	12	B <sup>1</sup> Σ <sub>u</sub> <sup>+</sup> 40,371	Gas	J. Phys. Chem. A. <b>113</b> , 11099 (2009)
		A <sup>1</sup> Δ <sub>u</sub> 25,046		Phys. Chem. Chem. Phys. <b>5</b> , 4772 (2003)
	14	B <sup>1</sup> Σ <sub>u</sub> <sup>+</sup> 37,667	Gas	J. Phys. Chem. A. <b>113</b> , 11099 (2009)
		A <sup>1</sup> Δ <sub>u</sub> 23,501		Phys. Chem. Chem. Phys. <b>5</b> , 4772 (2003)
	16	B <sup>1</sup> Σ <sub>u</sub> <sup>+</sup> 35,541	Gas	J. Chem. Phys. <b>114</b> , 2208 (2001)
	18	B <sup>1</sup> Σ <sub>u</sub> <sup>+</sup> 33,784	Gas	J. Chem. Phys. <b>114</b> , 2208 (2001)
	20	B <sup>1</sup> Σ <sub>u</sub> <sup>+</sup> 32,378	Gas	J. Chem. Phys. <b>114</b> , 2208 (2001)
	22	B <sup>1</sup> Σ <sub>u</sub> <sup>+</sup> 31,184	Gas	J. Chem. Phys. <b>114</b> , 2208 (2001)
	24	B <sup>1</sup> Σ <sub>u</sub> <sup>+</sup> 30,193	Gas	J. Chem. Phys. <b>114</b> , 2208 (2001)
	26	B <sup>1</sup> Σ <sub>u</sub> <sup>+</sup> 29,368	Gas	J. Chem. Phys. <b>114</b> , 2208 (2001)
<i>n</i> odd: <sup>3</sup> Σ <sub>u</sub> <sup>-</sup> ← X <sup>3</sup> Σ <sub>g</sub> <sup>-</sup>	3	C <sub>3</sub> H <sub>2</sub> X <sup>1</sup> A <sub>1</sub>	Gas	J. Mol. Spectrosc. <b>237</b> , 70 (2006)

(continued)

Table 2. Continued.

Species/electronic ground state	$n$	Origin (cm <sup>-1</sup> )	Phase	Reference
		A 13885–16389		
		B 16161–24802		
		C 39051–47156		
	5	A <sup>3</sup> Σ <sub>u</sub> <sup>-</sup> 23,033	Ne matrix	J. Chem. Phys. <b>103</b> , 8805 (1995)
	7	A <sup>3</sup> Σ <sub>u</sub> <sup>-</sup> 19,818	Gas	J. Chem. Phys. <b>119</b> , 814 (2003)
	9	A <sup>3</sup> Σ <sub>u</sub> <sup>-</sup> 17,185	Gas	J. Chem. Phys. <b>119</b> , 814 (2003)
	11	A <sup>3</sup> Σ <sub>u</sub> <sup>-</sup> 15,293	Gas	J. Chem. Phys. <b>119</b> , 814 (2003)
	13	A <sup>3</sup> Σ <sub>u</sub> <sup>-</sup> 13,916	Gas	J. Chem. Phys. <b>119</b> , 814 (2003)
		B <sup>3</sup> Σ <sub>u</sub> <sup>-</sup> 35,473		
	19	B <sup>3</sup> Σ <sub>u</sub> <sup>-</sup> 29,481	Gas	J. Chem. Phys. <b>119</b> , 814 (2003)
HC <sub><i>n</i></sub> H <sup>+</sup>	2	A <sup>2</sup> A 39,109	Gas	Chem. Phys. Lett. <b>174</b> , 109 (1990)
$n$ even X <sup>2</sup> Π		B <sup>2</sup> Σ <sup>+</sup> 56,380		
$n$ odd X <sup>2</sup> Π		C <sup>2</sup> Σ <sup>+</sup> 92,460		
	4	A <sup>2</sup> Π 19,722	Gas	Can. J. Phys. <b>34</b> , 1046 (1956)
		A <sup>2</sup> Π 19,645	Ne matrix	Chem. Phys. <b>189</b> , 335 (1994)
		C <sup>2</sup> Π 29,753		J. Phys. Chem. <b>111</b> , 1831 (2007)
	5	A <sup>2</sup> Π 20,026	Ne matrix	Chem. Phys. <b>189</b> , 335 (1994)
	6	A <sup>2</sup> Π 16,654	Gas	J. Chem. Phys. <b>89</b> , 6081 (1988)
		A <sup>2</sup> Π 16,535	Ne matrix	Chem. Phys. <b>189</b> , 335 (1994)
		C <sup>2</sup> Π 23,939		J. Phys. Chem. <b>111</b> , 1831 (2007)
	7	A <sup>2</sup> Π 16,676	Ne matrix	Chem. Phys. <b>189</b> , 335 (1994)
	8	A <sup>2</sup> Π 14,143	Gas	J. Chem. Phys. <b>110</b> , 296 (1999)
		A <sup>2</sup> Π 14,017	Ne matrix	Chem. Phys. <b>189</b> , 335 (1994)
		B <sup>2</sup> Π 17,590		J. Phys. Chem. <b>111</b> , 1831 (2007)
		C <sup>2</sup> Π 20,194		
		E <sup>2</sup> Π 41,254		
	9	A <sup>2</sup> Π 14,387	Ne matrix	Chem. Phys. <b>189</b> , 335 (1994)
	10	A <sup>2</sup> Π 12,260	Gas	Chem. Phys. Lett. <b>313</b> , 171 (1999)
		A <sup>2</sup> Π 12,143	Ne matrix	Chem. Phys. <b>189</b> , 335 (1994)
		B <sup>2</sup> Π 15,404		J. Phys. Chem. <b>111</b> , 1831 (2007)
		C <sup>2</sup> Π 17,627		
		E <sup>2</sup> Π 36,284		
	11	A <sup>2</sup> Π 12,674	Ne matrix	Chem. Phys. <b>189</b> , 335 (1994)
	12	A <sup>2</sup> Π 10,703	Ne matrix	Chem. Phys. <b>189</b> , 335 (1994)
		B <sup>2</sup> Π 13,708		J. Phys. Chem. <b>111</b> , 1831 (2007)
		D <sup>2</sup> Π 18,570		
		E <sup>2</sup> Π 32,382		
	13	A <sup>2</sup> Π 11,418	Ne matrix	Chem. Phys. <b>189</b> , 335 (1994)
	14	A <sup>2</sup> Π 9548	Ne matrix	Chem. Phys. <b>189</b> , 335 (1994)
		B <sup>2</sup> Π 12,447		J. Phys. Chem, 2007, 111, 1831
		D <sup>2</sup> Π 16,399		
		E <sup>2</sup> Π 31,095		
	15	A <sup>2</sup> Π 10,422	Ne matrix	Chem. Phys. <b>189</b> , 335 (1994)
	16	A <sup>2</sup> Π 8620	Ne matrix	Chem. Phys. <b>189</b> , 335 (1994)

(continued)

Table 2. Continued.

Species/electronic ground state	$n$	Origin (cm <sup>-1</sup> )	Phase	Reference
HC <sub><i>n</i></sub> H <sup>-</sup> 2Π (1,2,3) ← X 2Π	12	2Π 12,814	Ne matrix	J. Phys. Chem. A. <b>102</b> , 9785 (1998)
		2Π 13,990		
		2Π 31,506		
	14	2Π 11,633	Ne matrix	J. Phys. Chem. A. <b>102</b> , 9785 (1998)
		2Π 12,900		
		2Π 29,070		
	16	2Π 10,540	Ne matrix	J. Phys. Chem. A. <b>102</b> , 9785 (1998)
		2Π 12,522		
		2Π 26,940		
	18	2Π 9626	Ne matrix	J. Phys. Chem. A. <b>102</b> , 9785 (1998)
		2Π 10,506		
		2Π 27,248		
20	2Π 8809	Ne matrix	J. Phys. Chem. A. <b>102</b> , 9785 (1998)	
	2Π 9887			
	2Π 23,635			
22	2Π 8139	Ne matrix	J. Phys. Chem. A. <b>102</b> , 9785 (1998)	
	2Π 9393			
	2Π 22,282			
24	2Π 7545	Ne matrix	J. Phys. Chem. A. <b>102</b> , 9785 (1998)	
	2Π 21,218			
Odd $n$ C <sub>3</sub> H <sub>2</sub> <sup>-</sup> X <sup>2</sup> B <sub>1</sub>	3	A <sup>2</sup> A <sub>1</sub> 14,294	Gas	J. Chem. Phys. <b>105</b> , 10696 (1996) Phys. Chem. Chem. Phys. <b>3</b> , 4674 (2001)
		B <sup>2</sup> A'' 18,244		
		C <sup>2</sup> A <sub>1</sub> 20,055		
		D <sup>2</sup> B <sub>1</sub> 27,754		

support a DBS; instead it arises from the low-lying A 2Π state which is known to have a dipole moment around 4–5 D [45].

For the cationic species, electronic absorption spectra have been observed only for C<sub>6</sub>H<sup>+</sup> and C<sub>8</sub>H<sup>+</sup>. These have 3Σ<sup>-</sup> ground states and are isoelectronic with the corresponding neutral carbon chains. The electronic spectra have been assigned to the linear 3Σ<sup>-</sup> ← X 3Σ<sup>-</sup> transition and are found to lie in the vicinity of the corresponding transitions in C<sub>6</sub> and C<sub>8</sub>. In the case of C<sub>8</sub>H<sup>+</sup> additional weak transitions towards the UV have been observed as well.

### 3.2. HC<sub>*n*</sub>H, HC<sub>*n*</sub>H<sup>+</sup> and HC<sub>*n*</sub>H<sup>-</sup>

The even-membered neutral polyacetylenes are closed shell species with 1Σ<sub>g</sub><sup>+</sup> (π<sub>g</sub><sup>4</sup>/π<sub>u</sub><sup>4</sup>) ground state like the corresponding isoelectronic C<sub>*n*</sub>H<sup>-</sup> chains. The smaller members, HC<sub>4</sub>H and HC<sub>6</sub>H have been detected in the ISM [46–48]. π ← π excitation in these systems gives rise to the lowest excited states of 1,3Σ<sub>u</sub><sup>+</sup>, 1,3Σ<sub>u</sub><sup>-</sup> and 1,3Δ<sub>u</sub> symmetry. The intense dipole allowed B 1Σ<sub>g</sub><sup>+</sup> ← X 1Σ<sub>g</sub><sup>+</sup> transition has been observed in the gas phase for

$n=2, 4, \dots, 26$ , in neon matrices for  $n=12, 14, \dots, 24$  and in solution for  $n=10, 12, \dots, 20-24$ . The REMPI spectra of the above band system for  $n=16-26$  are displayed in Figure 3. The bands lie in the UV and are dominated by progressions in the acetylenic stretching mode in the excited state; the frequency decreasing with the chain size. The transitions also exhibit intrinsic lifetime broadening which has been attributed to fast intramolecular relaxation processes. Density functional theory (DFT) calculations predict a significant single–triple bond length alternation in the ground state indicating an acetylenic structure [49]. This is expected to be less pronounced in the excited state resulting in a cumulene-like structure. The dipole forbidden transitions to the  ${}^1\Delta_u$  and  ${}^1\Sigma_u^-$  states become allowed in the even chains via Herzberg–Teller vibronic coupling. These have been observed for  $n=4-14$  by REMPI.

The odd-membered polyacetylenes are open shell species with  ${}^3\Sigma_g^-$  ground states arising from  $\pi_g^4 \pi_u^2 / \pi_u^4 \pi_g^4$  electronic configuration.  $\pi \leftarrow \pi$  excitation from a fully occupied inner orbital to the HOMO and that from HOMO to lowest unoccupied molecular orbital (LUMO) gives two triplet excited states designated as A  ${}^3\Sigma_u^-$  and B  ${}^3\Sigma_u^-$ . The A  ${}^3\Sigma_u^- \leftarrow X {}^3\Sigma_g^-$  transition which lies in the visible has been observed for  $\text{HC}_n\text{H}$ ,  $n=5, 7, \dots, 15$  in neon matrices and for  $n=7, 9, 11$  and 13 in the gas phase. Calculations predict that the more intense B  ${}^3\Sigma_u^- \leftarrow X {}^3\Sigma_g^-$  transition begins in the VUV for  $\text{HC}_5\text{H}$  (170 nm) [50] and extends into the UV for  $\text{HC}_9\text{H}$  (210 nm) [51]. The B  ${}^3\Sigma_u^-$  state

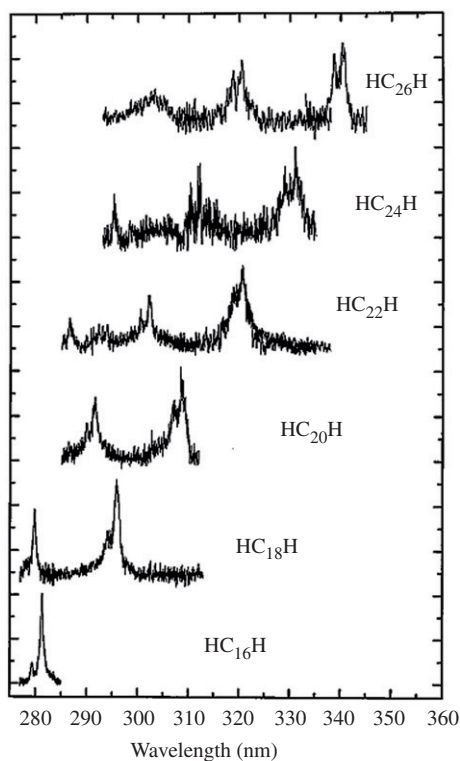


Figure 3. Resonant two-colour, two-photon ionisation spectra of the  ${}^1\Sigma_u^+ \leftarrow X {}^1\Sigma_g^+$  transition of polyynes  $\text{HC}_{2n}\text{H}$  ( $n=8-13$ ).

has been observed for HC<sub>13</sub>H (280 nm) and HC<sub>19</sub>H (340 nm) by REMPI. Oscillator strengths ( $f$ ) for the two transitions for  $n = 5, 7, \dots, 19$  were calculated at the CASSCF level of theory [52]. The  $f$ -value of the A  $\leftarrow$  X transition appears to first increase up to  $n = 7$  and then decrease with the chain size while for the B  $\leftarrow$  X transition  $f$  gradually increases, converging at higher  $n$ . This behaviour has been explained in terms of configuration mixing between the A and B states, beginning with nearly equal contribution from the two for smaller chains and gradually taken over by the B state for larger systems.

The even/odd numbered polyacetylene cations have  $^2\Pi$  ground states arising from  $\pi^3/\pi^1$  configuration. Electronic absorption spectra of the (A, B, C, D and E)  $^2\Pi \leftarrow X$   $^2\Pi$  transitions have been observed in neon matrices for  $n = 2, 4, \dots, 16$ . The A  $^2\Pi \leftarrow X$   $^2\Pi$  systems have also been studied in the gas phase for  $n = 2, 4$  by laser induced fluorescence (LIF),  $n = 6, 8$  and  $10$  by CRDS and  $n = 4, 6$  and  $8$  in a radio frequency trap by two-colour excitation–dissociation approach. For the odd chains  $n = 5, 7, \dots, 15$  another set of matrix absorptions have been tentatively assigned to the A  $^2\Pi \leftarrow X$   $^2\Pi$  transition. Like the cations, the polyacetylene anions have  $^2\Pi$  ground states as well. Three excited electronic states, all assigned to  $^2\Pi$  symmetry have been observed in 6 K neon matrices for even chains,  $n = 12, 14, \dots, 24$ . The first two  $^2\Pi \leftarrow X$   $^2\Pi$  transitions overlap in the IR and show progressions in the acetylenic stretching mode while the third  $^2\Pi \leftarrow X$   $^2\Pi$  transition lies in the UV. In the odd chains, linear C<sub>3</sub>H<sub>2</sub><sup>-</sup> isomer has been investigated by resonant two-photon detachment spectroscopy. The lowest excited state is a DBS while three higher excited states above the photodetachment threshold characterised as Feshbach states were observed.

## 4. Nitrogen and phosphorous terminated carbon chains

### 4.1. C<sub>n</sub>N, C<sub>n</sub>N<sup>+</sup> and C<sub>n</sub>N<sup>-</sup>

Among the radicals of the type C<sub>n</sub>N, the odd-membered species CN, C<sub>3</sub>N and C<sub>5</sub>N have been detected in the ISM [53,54]. These have been shown to have a  $^2\Sigma^+$  ground state arising from a  $\pi^4\sigma^1$  configuration. The higher members  $n \geq 9$  exhibit a  $^2\Pi$  ground state with a  $\pi^4\pi^1$  configuration similar to the even-membered chains [55,56]. As evident from Table 3, Panel A, experimental information on the excited electronic states of the even chains is sparse with the exception of C<sub>2</sub>N whose absorption spectrum was obtained in the flash photolysis of diazoacetonitrile. Recently, the LIF spectrum of the B  $^2\Pi \leftarrow X$   $^2\Sigma^+$  band system of C<sub>3</sub>N was measured and a Renner–Teller (RT) analysis of the B state was presented. Rotational analysis confirmed  $^2\Pi_{3/2}$  as the lower component indicating that the excited state is inverted with a  $\pi^3$  configuration. Figure 4 depicts the rotationally resolved LIF spectrum of the B  $^2\Pi_{3/2} \leftarrow X$   $^2\Sigma^+$  component of the origin band. The C<sub>2n-1</sub>N chains are isoelectronic with the C<sub>2n</sub>H species and like the latter have two low-lying  $^2\Sigma^+$  and  $^2\Pi$  states. In the case of C<sub>3</sub>N, dispersed fluorescence measurements from the B state have shown that the X  $^2\Sigma^+ - A$   $^2\Pi$  spacing is around 1840 cm<sup>-1</sup>. Theoretical studies indicate that the energy difference between these states decreases with the chain length leading to a change in the ground state symmetry from  $^2\Sigma^+$  to  $^2\Pi$  at C<sub>7</sub>N [55]. Electronic absorption spectra of the longer chains C<sub>2n-1</sub>N ( $n = 3-7$ ) have been recorded in 6 K neon matrices. These have been assigned to B  $^2\Pi \leftarrow X$   $^2\Pi$  for  $n = 5-7$  and B  $^2\Pi \leftarrow X$   $^2\Sigma^+$  for  $n = 3$  and  $4$ . Based on the similarities in the spectral profile and vibrational pattern, it has been proposed that C<sub>7</sub>N like C<sub>5</sub>N has a  $^2\Sigma^+$  ground state.



Table 3. N terminated carbon chains.

Species/electronic ground state	$n$	Origin (cm <sup>-1</sup> )	Phase	Reference
<i>Panel A: C<sub>n</sub>N species</i>	2	<b>CNC isomer</b>		
C <sub>n</sub> N		A <sup>2</sup> Δ 21,2534	Ne matrix	Helv. Chim. Acta. <b>84</b> , 1432 (2001)
$n$ even X <sup>2</sup> Π		B <sup>2</sup> Σ <sup>-</sup> 22,437		
$n$ odd X <sup>2</sup> Σ <sup>+</sup>		C <sup>2</sup> Σ <sup>+</sup> 26,652		
		D 27,330		
		A 21,259	Gas	Can. J. Phys. <b>44</b> , 353 (1996)
		B 22,413		
		C 26,662		
<b>CNC isomer</b>				
A <sup>2</sup> Δ <sub>u</sub> 30,303			Ne matrix	Helv. Chim. Acta. <b>84</b> , 1432 (2001)
B <sup>2</sup> Σ <sub>u</sub> <sup>-</sup> 34,662				
A 30,338			Gas	Can. J. Phys. <b>44</b> , 353 (1996)
B 34,802				
A <sup>2</sup> Π 1844	3		Gas	J. Chem. Phys. <b>127</b> , 184304 (2007)
B <sup>2</sup> Π 27,929				
B <sup>2</sup> Π 21,259	5		Ne matrix	J. Chem. Phys. <b>110</b> , 1492 (1999)
B <sup>2</sup> Π 17,185	7		Ne matrix	J. Chem. Phys. <b>110</b> , 1492 (1999)
B <sup>2</sup> Π 14,802	9		Ne matrix	J. Chem. Phys. <b>110</b> , 1492 (1999)
B <sup>2</sup> Π 13,207	11		Ne matrix	J. Chem. Phys. <b>110</b> , 1492 (1999)
B <sup>2</sup> Π 12,085	13		Ne matrix	J. Chem. Phys. <b>110</b> , 1492 (1999)
A <sup>1</sup> Σ <sup>+</sup> 36,496	7		Ne matrix	J. Chem. Phys. <b>110</b> , 1492 (1999)
A <sup>1</sup> Σ <sup>+</sup> 30,656	9		Ne matrix	J. Chem. Phys. <b>110</b> , 1492 (1999)
A <sup>1</sup> Σ <sup>+</sup> 26,295	11		Ne matrix	J. Chem. Phys. <b>110</b> , 1492 (1999)
A <sup>1</sup> Σ <sup>+</sup> 22,619	13		Ne matrix	J. Chem. Phys. <b>110</b> , 1492 (1999)
C <sub>n</sub> N <sup>+</sup> A <sup>1</sup> Σ <sup>+</sup> ← X <sup>1</sup> Σ <sup>+</sup>	2	CCN <sup>+</sup> 21,637	Ne matrix	Helv. Chim. Acta. <b>84</b> , 1432 (2001)
$n$ even X <sup>3</sup> Σ <sup>-</sup>		CNC <sup>+</sup> 30,703		
$n$ odd X <sup>1</sup> Σ <sup>+</sup>				
C <sub>n</sub> N <sup>+</sup> A <sup>1</sup> Π ← X <sup>1</sup> Σ <sup>+</sup>				

Panel B:  $NC_nN$  species

$NC_nN$	2	$^1\Sigma_u^+$ 75,757 $^1\Pi_u$ 60,422	UV absorption	J. Mol. Spectrosc. <b>30</b> , 162 (1969) J. Mol. Spectrosc. <b>41</b> , 20 (1972)
$n$ even $X$ $^1\Sigma_g^+$		$^1\Delta_u$ 48,309		J. Mol. Spectrosc. <b>39</b> , 393 (1971)
$n$ odd $X$ $^3\Sigma_g^-$		$^1\Sigma_u^+$ 45,454 $^3\Delta_u$ 40,000 $^3\Sigma_u^-$ 33,333		Proc. Phys. Soc. Lond. <b>82</b> , 335 (1963)
	3	A 27,850 B 37,310	Ar matrix	Spectrochim. Acta. <b>50A</b> , 209 (1994)
	4	$^1\Delta_u$ 37,313 $^1\Sigma_u^-$ 35,714	UV absorption	Spectrochim Acta. <b>12</b> , 521 (1958)
	5	$^3\Sigma_u^-$ 22,836 $^3\Sigma_u^-$ 22,763	Ne matrix	Chem. Phys. Lett. <b>351</b> , 85 (2002) Chem. Phys. Lett. <b>345</b> , 89 (2001)
	6	$^1\Sigma_u^+$ 32,258	UV absorption	J. Chem. Phys. <b>52</b> , 5184 (1970)
	8	$^1\Sigma_u^+$ 43,103	Solution	Chem. Eur. J. <b>3</b> , 1105 (1997)
	10	$^1\Sigma_u^+$ 38,610	Solution	Chem. Eur. J. <b>3</b> , 1105 (1997)
	12	$^1\Sigma_u^+$ 35,460	Solution	Chem. Eur. J. <b>3</b> , 1105 (1997)
	14	$^1\Sigma_u^+$ 33,003	Solution	Chem. Eur. J. <b>3</b> , 1105 (1997)
	16	$^1\Sigma_u^+$ 31,056	Solution	Chem. Eur. J. <b>3</b> , 1105 (1997)
	18	$^1\Sigma_u^+$ 34,916 $^1\Sigma_u^+$ 29,498	Gas	Chem. Phys. Lett. <b>393</b> , 225 (2004) Chem. Eur. J. <b>3</b> , 1105 (1997)
$NC_nN^+$	2	A $^2\Sigma_u^+$ 9340 B $^2\Sigma_u^+$ 11,290 C $^2\Pi_u$ 17,310 D $^2\Pi_u$ 33522	Gas	J. Chem. Phys. <b>123</b> , 144302 (2005) J. Chem. Phys. <b>129</b> , 134315 (2008)
$(n$ even, $X$ $^2\Pi_{g(u)}$ )		$A^2\Pi_g$ 16,780		
$(n$ odd, $X$ $^2\Pi_{g(u)}$ )		$B^2\Sigma_g^+$ 17,430 C $^2\Sigma_u^+$ 18,720 D $^2\Pi_u$ 25,500		
$NC_nN^+$	4	$A^2\Pi_u$ 15,246	Gas	J. Phys. Chem. <b>86</b> , 514 (1982)
$n$ even $X$ $^2\Pi_{g(u)}$	6		Gas	J. Chem. Phys. <b>71</b> , 3125 (1979)

(continued)

Table 3. Continued.

Species/electronic ground state	$n$	Origin ( $\text{cm}^{-1}$ )	Phase	Reference
$n$ odd $X^2\Pi_{g/u}$		$B^2\Sigma_g^+, C^2\Sigma_u^+$ 22,190		J. Chem. Phys. <b>116</b> , 924 (2002)
		$D^2\Pi_g$ 25,580		
		$E^2\Pi_u$ 30,420		
		$A^2\Pi_u$ 15173		
	7	15,897	Ne matrix	J. Chem. Phys. <b>102</b> , 1510 (1995)
	8	$A^2\Pi_g$ 13,489	Ne matrix	J. Chem. Phys. <b>102</b> , 1510 (1995)
	9	14,024	Ne matrix	J. Chem. Phys. <b>102</b> , 1510 (1995)
	10	$A^2\Pi_u$ 12,029	Ne matrix	J. Chem. Phys. <b>102</b> , 1510 (1995)
	11	12,587	Ne matrix	J. Chem. Phys. <b>102</b> , 1510 (1995)
	12	$A^2\Pi_g$ 10,831	Ne matrix	J. Chem. Phys. <b>102</b> , 1510 (1995)
<i>Panel C: HC<sub>n</sub>N species</i>				
$HC_nN$	2	a 4155	Gas	J. Chem. Phys. <b>117</b> , 4323 (2002)
$n$ odd $X^1\Sigma^+$		$A^3\Sigma^-$ 41,667–29,412	Ar matrix	J. Chem. Phys. <b>66</b> , 4334 (1977)
$n$ even $X^3\Sigma^-$	3	$A^1A''$ 38,485	Gas	J. Mol. Spectrosc. <b>19</b> , 155 (1966)
		$B^1\Delta/\Sigma^-$ 44,221	Gas	J. Mol. Spectrosc. <b>19</b> , 178 (1966)
	6	$A^3\Sigma^-$ 21,209	Gas	J. Chem. Phys. <b>114</b> , 7918 (2001)
		$A^3\Sigma^-$ 21,181	Ne matrix	
$HC_nN^+$	3	A 15,650	Gas	Proc. Roy. Soc. (London). <b>A308</b> , 19 (1968)
$n$ odd $^2\Pi$		B 19,600		
$n$ even $^2\Pi$		C 48,570		
		B 19,374		
	5	$A^2\Pi$ 17,178	Ne matrix	J. Phys. Chem. <b>89</b> , 3190 (1985)
		$B^2\Sigma^+$ 21860	Gas	J. Chem. Soc. Faraday Trans. <b>276</b> , 676 (1980)
		$C^2\Pi$ 27350		J. Chem. Phys. <b>111</b> , 9600 (1999)

HC <sub>n</sub> N <sup>-</sup> <i>n</i> odd <sup>2</sup> Π <i>n</i> even <sup>2</sup> Π	7	A <sup>2</sup> Π 17,130 A <sup>2</sup> Π 14,925 A <sup>2</sup> Π 14,845 15,217 A <sup>2</sup> Π 12,978 13,395 A <sup>2</sup> Π 11,468 12,026 A <sup>2</sup> Π 10,276 2Π 13,466 2Π 14,196 2Π 32,895 2Π 12,038 2Π 12,692 2Π 29,762 2Π 10,811 2Π 11,500 2Π 27,525 2Π 9745 2Π 10,611 2Π 25,562 2Π 8864 2Π 9950 2Π 25,562 2Π 8108 2Π 9413 2Π 22,437 2Π 7483 2Π 28,952	Ne matrix Gas Ne matrix Ne matrix Ne matrix Ne matrix Ne matrix Ne matrix Ne matrix Ne matrix Ne matrix	J. Chem. Phys. <b>102</b> , 1510 (1995) J. Chem. Phys. <b>112</b> , 8899 (2000) J. Chem. Phys. <b>102</b> , 1510 (1995) J. Chem. Phys. <b>102</b> , 1510 (1995) J. Chem. Phys. <b>102</b> , 1510 (1995) J. Chem. Phys. <b>102</b> , 1510 (1995) J. Chem. Phys. <b>102</b> , 1510 (1995) J. Chem. Phys. <b>102</b> , 1510 (1995) J. Chem. Phys. <b>102</b> , 1510 (1995) J. Chem. Phys. <b>102</b> , 1510 (1995) J. Phys. Chem. A. <b>102</b> , 9785 (1998)
	13		Ne matrix	J. Phys. Chem. A. <b>102</b> , 9785 (1998)
	15		Ne matrix	J. Phys. Chem. A. <b>102</b> , 9785 (1998)
	17		Ne matrix	J. Phys. Chem. A. <b>102</b> , 9785 (1998)
	19		Ne matrix	J. Phys. Chem. A. <b>102</b> , 9785 (1998)
	21		Ne matrix	J. Phys. Chem. A. <b>102</b> , 9785 (1998)
	23		Ne matrix	J. Phys. Chem. A. <b>102</b> , 9785 (1998)

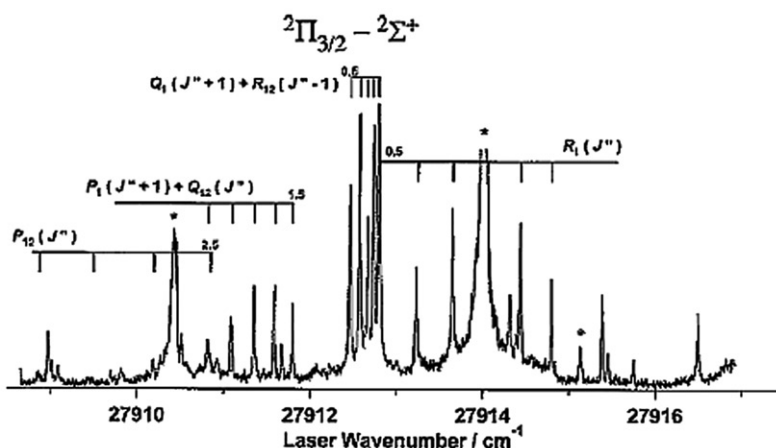


Figure 4. High-resolution laser induced fluorescence spectrum of the  ${}^2\Pi_{3/2} \leftarrow X {}^2\Sigma^+$  component of the origin band of the  $C_3N$ . The strong lines marked by asterisks are those of  $CN B {}^2\Sigma^+ \leftarrow X {}^2\Sigma^+$ . Reproduced with permission from *J. Chem. Phys.* **127**, 184304 (2007).

Table 4. P terminated carbon chains:  $C_nP$  species.

Species/Electronic ground state	$n$	Origin ( $\text{cm}^{-1}$ )	Phase	Reference
$C_nP$ ( $n \geq 2$ , $X {}^2\Pi$ )	2	${}^2\Delta$ 15,898	Gas	<i>J. Chem. Phys.</i> <b>128</b> , 244311 (2008)

The ground state symmetries of even and odd-membered anions are known to be  ${}^1\Sigma^+$  ( $\pi^4\sigma^2$ ) and  ${}^3\Sigma^-$  ( $\pi^4\pi^2$ ), respectively. Recently,  $C_3N^-$  and  $C_5N^-$  were detected in the circumstellar envelope of IRC+10216 [57,58]. Spectra of the  $A {}^1\Sigma^+ \leftarrow X {}^1\Sigma^+$  transition of odd chains,  $C_{2n-1}N^-$  ( $n=4-7$ ) have been measured in neon matrices. For the even chains,  $C_{2n}N^-$  ( $n=1-3$ ), high-resolution photoelectron spectra have been recorded [59]. The results confirm that  $C_2N^-$ ,  $C_4N^-$  and  $C_6N^-$  all have  ${}^3\Sigma^-$  ground states. In addition, the corresponding neutrals have  ${}^2\Pi_r$  ground states with a spin-orbit splitting of 38, 29 and  $30 \text{ cm}^{-1}$ , respectively. In the case of the cations, removal of one electron from either  $\sigma$  or  $\pi$  orbital corresponding to the odd and even neutral ground states results in a  ${}^1\Sigma^+$  symmetry. Experimental data are only available for  $C_2N^+$  where the IR and UV spectra for  $CCN^+$  and  $CNC^+$  in the  $X {}^1\Sigma^+$  and  $A {}^1\Pi$  states were measured in neon matrices (Table 3).

#### 4.2. $HC_nN$ , $HC_nN^+$ and $HC_nN^-$

Cyanopolynes of the type  $HC_{2n-1}N$  are closed shell species with  $\pi^4\sigma^2$  electron configuration and  ${}^1\Sigma^+$  ground states. Rotational spectra up to  $HC_{17}N$  [60] have been measured by microwave spectroscopy and chains up to  $HC_{11}N$  have been detected in space [61]. These studies confirm that the odd cyanopolynes possess linear ground state geometry and show significant single-triple bond length alternation, much like the

isoelectronic polyacetylenes,  $\text{HC}_{2n}\text{H}$ . In contrast, the ground state geometry in the even chains remains ambiguous due to competition between closely lying linear and ring chain isomers. For example, calculations [62,63] predict two stable isomers for  $\text{HC}_4\text{N}$ , the linear,  $^3\Sigma^-$  and the  $^1\text{A}'$  ring chain forms and three isomers for  $\text{HC}_6\text{N}$ , a triplet linear and two singlet ring chain structures. In  $\text{HC}_4\text{N}$ , the  $^1\text{A}'$  is more stable by  $20\text{ kJ mol}^{-1}$  while in  $\text{HC}_6\text{N}$ , the linear  $^3\Sigma^-$  is lower in energy by 10 and  $16.7\text{ kJ mol}^{-1}$ , respectively, from the two  $^1\text{A}'$  isomers. In two different mm wave studies, the ring chain structure has been observed for both species [64], while the linear isomer has been detected for  $\text{HC}_4\text{N}$  [65].

In the case of  $\text{HC}_2\text{N}$ , an electronic band system was observed in the 240–340 nm region in an argon matrix and tentatively assigned to the  $\text{A}^3\Sigma^- \leftarrow \text{X}^3\Sigma^-$  transition. In a combined matrix and CRD study, the origin band of the corresponding transition in  $\text{HC}_6\text{N}$  was detected around 471 nm and rotationally resolved. These transitions involve  $\pi^* \leftarrow \pi$  excitation and show progressions in the acetylenic stretching mode. Similar features have been observed in the electronic spectra of  $\text{NC}_5\text{N}$  and  $\text{HC}_7\text{H}$  which are isoelectronic with  $\text{HC}_6\text{N}$ . Gas phase absorption spectrum of  $\text{HC}_3\text{N}$  revealed two-band systems in the UV, one corresponding to the linear–bent transition,  $\text{A}^1\text{A}'' \leftarrow \text{X}^1\Sigma^+$ , similar to the 230 nm band system in acetylene and the other to a  $^1\Delta$  or  $^1\Sigma^+$  excited state, electronically forbidden, but made allowed by  $\pi$ -type vibrations.

The odd and even numbered cyanopolyne anions have  $^2\Pi$  ground states with a  $\pi^4\pi^1$  or  $\pi^4\pi^3$  configuration. Three band systems (A,B,C)  $^2\Pi \leftarrow \text{X}^2\Pi$  were observed for the  $\text{HC}_{2n-1}\text{N}^-$  ( $n=7-14$ ) chains in neon matrices with the A and B systems overlapping in the near IR and the C state in the UV. The A and B states show strong origin bands which are separated by  $730\text{ cm}^{-1}$  in  $\text{HC}_{11}\text{N}^-$  and move farther apart as the chain increases approaching  $1305\text{ cm}^{-1}$  in  $\text{HC}_{21}\text{N}^-$ . The B state gradually loses intensity with the chain length – this has been attributed to the decrease in the magnitude of vibronic interaction between the states as they move apart. For the three excited states, the band origin shifts to the red with number of carbon atoms, characteristic of  $\pi \leftarrow \pi$  transitions. Among the cyanopolyacetylene cations,  $\text{HC}_{2n-1}\text{N}^+$  ( $n=2-7$ ) and  $\text{HC}_{2n}\text{N}^+$  ( $n=3-6$ ), electronic transitions have been observed in the gas phase and neon matrices. The origin bands of the  $\text{A}^2\Pi \leftarrow \text{X}^2\Pi$  systems in  $\text{HC}_5\text{N}^+$  and  $\text{HC}_7\text{N}^+$  have been rotationally resolved and analysed using frequency modulation absorption spectroscopy. Based on the rotational constants and molecular orbital consideration, it has been suggested that a reduction in the antibonding character of the C–C and bonding character of  $\text{C}\equiv\text{C}$  determine the structural changes upon ionisation to the X and A states of the cation. The ground  $\text{X}^2\Pi$  state results from the removal of one electron from the  $\pi$  orbital of the neutral which is bonding in the  $\text{C}\equiv\text{C}$  regions and antibonding in the C–C regions resulting in a net contraction of the molecule and therefore a higher  $B_0$  value than the neutral. In contrast, the  $\text{A}^2\Pi$  state involves loss of an electron from the  $\pi$  orbital which is bonding in the outer  $\text{C}\equiv\text{C}$  regions and slightly bonding in the C–C region, resulting in a net increase in the chain length and smaller  $B_0$  value than the neutral.

### 4.3. $\text{NC}_n\text{N}$ and $\text{NC}_n\text{N}^+$

The dicyanopolyynes have linear ground state structures with  $\text{D}_{\infty\text{h}}$  symmetry where the N atoms bond to the C atoms at the end of the chain. The even members,  $\text{NC}_{2n}\text{N}$  have  $\text{X}^1\Sigma_g^+$  ( $\pi_g^4/\pi_u^4$ ) ground states while the odd systems have  $\text{X}^3\Sigma_g^-$  ( $\pi_g^2/\pi_u^2$ ) symmetry. These chains

are thought of as resulting from a mixture of resonance structures with the even- and odd-numbered chains tending towards acetylenic and cummulenic forms, respectively [56,66]. UV absorption spectra have been measured for  $\text{NC}_{2n}\text{N}$ ,  $n=1-3, 8$  in the gas phase and  $n=4-9$  in solution. In the case of  $\text{NC}_{16}\text{N}$ , the  $\text{B } ^1\Sigma_u^+ \leftarrow \text{X } ^1\Sigma_g^+$  transition was studied using a REMPI technique. The location of the band origin and the spectral profile closely resembled that of the isoelectronic  $\text{HC}_{18}\text{H}$  indicating similar electronic structures. In addition, the ground and excited states  $\text{A } ^1\Sigma_u^+$ ,  $\text{A } ^1\Delta_u$  and  $\text{B } ^1\Sigma_u^+$  were investigated using time-dependent density functional theory. These studies suggest that in the  $\text{NC}_{2n}\text{N}$  series, like the  $\text{HC}_{2n}\text{H}$  chains, the single-triple bond length alternation is significant in the ground state and less pronounced in the excited state. Accordingly, the former assumes an acetylenic form whereas the latter are cummulenic in nature. For the odd numbered chains, the  $\text{A } ^3\Sigma_u^- \leftarrow \text{X } ^3\Sigma_g^-$  band system in  $\text{NC}_5\text{N}$  has been rotationally resolved using CRDS, while two excited states have been identified in argon matrices and assigned to  $\text{NC}_3\text{N}$ . The band origins and excited state symmetries are presented in Table 3, Panel B.

The even and odd dicyanopolyne cations have  $\text{X } ^2\Pi_{g/u}$  ground states. For the even-numbered chains, the  $\text{X } ^2\Pi_{g/u}$  state can be described by  $\sigma_u^2\sigma_g^2\pi_{g/u}^4\pi_{u/g}^3$  electron configuration. Promotion of an electron from the fully occupied inner  $\pi$  and  $\sigma$  orbitals to the HOMO gives  $^2\Pi_{g/u}$ ,  $^2\Sigma_g^+$  and  $^2\Sigma_u^+$  excited states. Electronic excitation from the  $\pi_{u/g}^3$  HOMO to  $\pi_{g/u}$  LUMO results in  $^2\Pi_{g/u}$  states which lie higher in energy; these have been tabulated in Table 3, Panel B. Recently, the  $\text{B } ^2\Sigma_u^+ \leftarrow \text{X } ^2\Pi_g$  absorption spectrum of  $\text{NC}_2\text{N}^+$  was measured in an ion trap using a two-colour, two-photon dissociation technique. Figure 5 depicts the absorption spectrum recorded at  $\sim 20$  K and  $0.3\text{ cm}^{-1}$  resolution, showing excitation of four normal modes;  $\nu_1(\sigma_g^+)$ ,  $\nu_2(\sigma_g^+)$ ,  $\nu_4(\pi_g)$  and  $\nu_5(\pi_u)$ . The  $\text{A } ^2\Pi_u \leftarrow \text{X } ^2\Pi_g$  transition of  $\text{NC}_6\text{N}^+$  has been rotationally analysed using frequency modulation absorption technique. Spectral perturbations characteristic of  $\Lambda$ -type doubling and an estimation of the second order spin-orbit splitting parameter of the  $\text{A } ^2\Pi_u$  state was used to identify and locate the unobserved  $^2\Sigma_g^+$ ,  $208\text{ cm}^{-1}$  below the A state. The first excited  $\text{A } ^2\Pi$  state has been identified in neon matrices for  $\text{NC}_{2n}\text{N}^+$ ,  $n=3-6$  along with unassigned states for  $\text{NC}_{2n-1}\text{N}^+$ ,  $n=4-6$ .

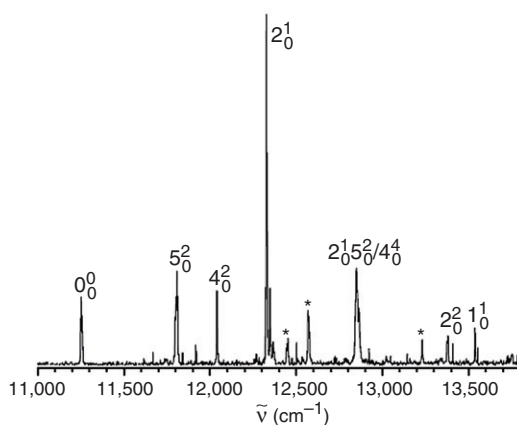


Figure 5. Resonant two-colour, two-photon spectrum of the  $\text{B } ^2\Sigma_u^+ \leftarrow \text{X } ^2\Pi_{3/2,g}$  transition in a 22-pole ion trap at 20 K recorded by monitoring  $\text{CN}^+$  as the photo-dissociation product.

#### 4.4. $C_nP$ , $C_nP^+$ and $C_nP^-$

In contrast to the nitrogen doped carbon chains, information on the electronic structure of the isovalent phosphorous terminated systems is scarce. The role of P containing molecules in the ISM attracted much attention following the detection of PN [67,68] and PC [69] in space.  $C_2P$  and  $HC_2P$  were predicted to be present in the warm dense clouds with the former being recently detected in IRC + 10216 [70]. Theoretical study of  $C_nP$ ,  $C_nP^+$  and  $C_nP^-$  ( $n=1-7$ ) [71] predict that these molecules assume a linear  $C_{\infty v}$  ground state geometry with the phosphorous atom located at the end of the carbon chain. With the exception of CP which has a  $^1\Sigma^+$  ground electronic state, the other members possess  $^2\Pi$  electronic ground state. The corresponding ions show odd/even alternation in the ground state symmetry; with  $^1\Sigma^+$  ( $n$  odd)/ $^3\Sigma^-$  ( $n$  even) for  $C_nP^-$  and  $^3\Sigma^-$  ( $n$  odd)/ $^1\Sigma^+$  ( $n$  even) for  $C_nP^+$ . Among the phosphorous terminated systems with  $n \geq 2$ ,  $C_2P$  is the only radical where the electronic structure has been experimentally investigated. Recently, a LIF spectrum of the  $^2\Delta \leftarrow X^2\Pi$  transition of  $C_2P$  (Table 4) was recorded and RT parameters for the ground and excited states were deduced. An allenic structure for the ground state has been inferred following rotational analyses of the two spin-orbit components of the origin band.

### 5. Oxygen terminated carbon chains

#### 5.1. $C_nO$ , $C_nO^+$ and $C_nO^-$

Carbon monoxide is the second most abundant molecular species in space following hydrogen. The next higher members,  $C_2O$  and  $C_3O$  have been detected in TMC-1 [72,73], thereby triggering the radio astronomical search for the longer chains in this family of polycarbon oxides. In terrestrial environment, the ground electronic states of  $C_nO$  ( $n=2-9$ ) have been investigated by microwave spectroscopy [74-76]. These studies have demonstrated that  $C_nO$ 's have linear ground state geometries with cumulenic bonds where ground state alternates between  $^3\Sigma^-$  ( $\pi^2$ ) for even and  $^1\Sigma^+$  ( $\pi^4$ ) for the odd chains. In addition, electron spin resonance spectra in low-temperature matrices measured for  $C_2O$ ,  $C_4O$  and  $C_6O$  support the above assignment. For the even chains, the  $\pi^2$  electron configuration also gives rise to  $^1\Sigma^+$  and  $^1\Delta$  states. In the microwave investigation on  $C_nO$  ( $n=2, 4, 6, 8$ ) [75], the spin-spin coupling parameter for the triplet state was deduced and was seen to increase rapidly with  $n$ . It was concluded that the dominant contribution was from the second-order spin-orbit coupling term which mixes the  $^3\Sigma^-$  and  $^1\Sigma^+$  states. As a result, the energy difference between these states tends to decrease with an increase in the chain length. The  $^1\Sigma^+$  state for  $n=2, 4$  and  $6$  has been predicted to lie around 1.10, 0.47 and 0.11 eV above the ground state. In an anion photodetachment study on  $C_2O^-$  [77], the  $^1\Sigma^+$  was identified and located around 1.01 eV above the triplet state. In addition, term energies of the  $^1\Delta$  and  $^3\Pi$  states were determined as 0.65 and 1.44 eV, respectively. The first optically allowed  $A^3\Pi \leftarrow X^3\Sigma^-$  transition has been observed by LIF (Table 5, Panel A) in which the upper state was found to be RT active; the ground and excited state molecular parameters were deduced from their rotational spectra.

The  $C_nO^\pm$  are expected to have linear ground state geometries with the even/odd-numbered cations and anions exhibiting  $^2\Pi_r$  ( $\pi^1$ )/ $^2\Pi_i$  ( $\pi^3$ ) and  $^2\Pi_i$  ( $\pi^3$ )/ $^2\Pi_r$  ( $\pi^1$ ) symmetries, respectively. There is a lack of experimental information on the electronic structure of the  $C_nO$  cations. In contrast, the smaller anions have been studied by matrix



isolation ( $C_2O^-$ ,  $C_4O^-$ ) and photoelectron ( $C_2O^-$ ,  $C_3O^-$  and  $C_4O^-$ ) spectroscopy [77–80]. The  $A \ ^2\Sigma^+ \leftarrow X \ ^2\Pi$  ( $C_2O^-$  and  $C_4O^-$ ) and the  $B \ ^2\Pi \leftarrow X \ ^2\Pi$  ( $C_4O^-$ ) transitions have been observed in neon matrices. A spin–orbit splitting of  $65 \text{ cm}^{-1}$  in  $X \ ^2\Pi_i$  of  $C_2O^-$  has been deduced from photoelectron studies [77]; in a separate investigation,  $C_3O^-$  was shown to have a bent equilibrium ground state [80].

### 5.2. $HC_nO$ , $HC_nO^+$ and $HC_nO^-$

The hydrogen substituted  $C_nO$ 's have received attention due to their role in combustion processes [81] and interstellar chemistry [82].  $HC_2O$  in particular has been recognised as a key intermediate in the oxidation of acetylene which ultimately results in the production of carbon dioxide. The geometrical structures of  $HC_nO$  ( $n=1-7$ ) have been characterised by studies in the mm and sub-mm regions [83–85]. The lower members with  $n=1-4$  are reported to have bent planar geometries with  $HC_3O$  and  $HC_4O$  being bent at the CCO subunit while  $HC_2O$  assumes a bent HCC structure. These radicals are subjected to strong RT interaction where the ground state alternates between  $^2A'$  for  $n$  odd and  $^2A''$  for  $n$  even which correlate with the  $\Pi$  state in the linear configuration. This energy difference has been estimated to be around  $540 \text{ cm}^{-1}$  for  $HC_2O$  and  $2560 \text{ cm}^{-1}$  for  $HC_3O$ . In contrast, the higher members with  $n=5, 6$  and  $7$  are found to have linear structures with  $^2\Pi$  ground states which alternate between  $^2\Pi_{3/2}$  for  $n$  even and  $^2\Pi_{1/2}$  for  $n$  odd.

In the case of  $HC_2O$ , the next higher state assigned to the linear  $B \ ^2\Pi_i$  has been observed by LIF spectroscopy. The spectrum revealed a progression in the  $\nu_3$ , CCO stretching mode and several other bands which were attributed to excitation of the RT active CCH and CCO bending modes in the excited state. In addition, the spin–orbit splitting in the  $B \ ^2\Pi_i$  state was deduced to be  $-46.73 \text{ cm}^{-1}$ . Experimental data on the ground and excited states of the corresponding anions and cations have not been reported thus far.

### 5.3. $OC_nO$

Among the polycarbon dioxides, electronic absorption spectra have been measured for the neutrals. In the linear configuration, the odd and even members of this group are expected to have  $^1\Sigma^+$  ( $\pi^4$ ) and  $^3\Sigma^-$  ( $\pi^2$ ) ground electronic states, respectively. The vacuum UV spectra of carbon suboxide,  $C_3O_2$  is dominated by an intense broad continuum around 159 nm, a band system around 178 nm assigned to the  $^1\Sigma_u^+ \leftarrow X \ ^1\Sigma_g^+$  transition and a weak continuum around 270 nm. UV absorption bands of the higher members have been observed in argon matrices and the wavelength positions are presented in Table 5, Panel C. Recently, the IR and UV-Vis spectra of  $OC_nO$  and  $C_nO$  ( $n=9, 11, \dots, 21$ ) trapped in low temperature matrices were experimentally and theoretically investigated. Electronic absorptions around 397 and 452 nm were tentatively assigned to  $C_{13}O$  and  $C_{15}O_2$ , respectively. The result from this study [86] also indicated that the polycarbon oxides possess strong electronic absorptions which lie in the vicinity of the corresponding bare carbon chains.

Table 5. O terminated carbon chains.

Species/electronic ground state	<i>n</i>	Origin (cm <sup>-1</sup> )	Phase	Reference
<i>Panel A: C<sub>n</sub>O species</i>				
C <sub><i>n</i></sub> O	2	a <sup>1</sup> Δ5242	Gas	J. Chem. Phys. <b>129</b> , 074312 (2008)
<i>n</i> even X <sup>3</sup> Σ <sup>-</sup>		b <sup>1</sup> Σ <sup>+</sup> 8416		J. Chem. Phys. <b>108</b> , 4070 (1998)
<i>n</i> odd X <sup>1</sup> Σ <sup>+</sup>		A <sup>3</sup> Π 11,614		Chem. Phys. <b>61</b> , 451 (1981)
		c <sup>1</sup> Π 22,390		
C <sub><i>n</i></sub> O <sup>-</sup>	2	A <sup>2</sup> Σ <sup>+</sup> 12234	Ne matrix	J. Phys. Chem. A. <b>102</b> , 3459 (1998)
<i>n</i> even X <sup>2</sup> Π	4	A <sup>2</sup> Σ <sup>+</sup> 11434	Ne matrix	J. Phys. Chem A. <b>105</b> , 4894 (2001)
<i>n</i> odd X <sup>2</sup> Π		B <sup>2</sup> Π 22610		
<i>Panel B: HC<sub>n</sub>O species</i>				
HC <sub><i>n</i></sub> O	2	B <sup>2</sup> Π 33,465	Gas	J. Chem. Phys. <b>110</b> , 6773 (1999)
<i>n</i> even X <sup>2</sup> A''				
<i>n</i> odd X <sup>2</sup> A'				
<i>Panel C: OC<sub>n</sub>O species</i>				
OC <sub><i>n</i></sub> O	3	<sup>1</sup> Σ <sup>+</sup> 56,179	Gas	J. Mol. Spectrosc. <b>21</b> , 42 (1996)
<i>n</i> even X <sup>3</sup> Σ <sup>-</sup>	4	47,200	Ar matrix	Angew. Chem. <b>102</b> , 920 (1990)
<i>n</i> odd X <sup>1</sup> Σ <sup>+</sup>	5	43,500	Ar matrix	Angew. Chem. <b>27</b> , 566 (1998)
	6	39,682	Ar matrix	Tetrahed. Lett. <b>32</b> , 4469 (1991)
	7	34,720	O <sub>2</sub> matrix	J. Phys. Chem. A. <b>110</b> , 12395 (2006)
	15	22,123	O <sub>2</sub> matrix	J. Mol. Spectrosc. <b>243</b> , 189 (2007)

## 6. Sulphur terminated carbon chains

### 6.1. C<sub>*n*</sub>S, C<sub>*n*</sub>S<sup>+</sup> and C<sub>*n*</sub>S<sup>-</sup>

The fractional abundance of sulphur in space is lower than oxygen; however, polycarbon sulphides C<sub>2</sub>S and C<sub>3</sub>S have been detected via radio astronomy in the dense molecular cloud TMC-1 [87], and are found to be present in greater abundance than the isovalent species, C<sub>2</sub>O and C<sub>3</sub>O. In addition, these have been identified in the outer envelopes of IRC+10216 along with C<sub>5</sub>S [88,89]. It has been postulated that the hydrogenated counterparts may be present in detectable abundances as well. The symmetrical species, SC<sub>*n*</sub>S however, lack a permanent dipole moment and cannot be detected via radio astronomy, but can be observed by their infrared or electronic absorptions.

The neutral C<sub>*n*</sub>S species show an alternation in the ground state symmetry. The even and odd numbered chains can be differentiated by <sup>3</sup>Σ<sup>-</sup> (π<sup>2</sup>) and <sup>1</sup>Σ<sup>+</sup> (π<sup>4</sup>) whereas the even/odd cations have <sup>2</sup>Π<sub>r</sub> (π<sup>1</sup>)/<sup>2</sup>Π<sub>i</sub> (π<sup>3</sup>) and the anions show <sup>2</sup>Π<sub>i</sub> (π<sup>3</sup>)/<sup>2</sup>Π<sub>r</sub> (π<sup>1</sup>) ground states. DFT calculations have been performed on neutrals up to *n*=20 and the ground state geometries have been compared with the bare carbon chains [90]. These studies indicate that the C<sub>*n*</sub>S molecules are linear up to *n*=17 beyond which there is a substantial decrease in the energy difference between the linear and cyclic forms. The latter is more stable by 24.3 kJ/mol for *n*=18, while *n*=19, 20 are linear. In contrast, the carbon species are predominantly linear up to *n*=10 while the larger members assume cyclic structures. This has been attributed to the divalent nature of sulphur as opposed to the tetravalent one of carbon. As a result, sulphur takes up the terminal position of a C-chain where it is doubly

bonded to the adjacent carbon, while in the cyclic structure; it lies outside the ring and forms two single bonds with two neighbouring carbon atoms. This preference to linearity has been observed in the anions and cations as well [91]. Linear structures have been predicted for anion chains up to  $n=16$  and  $n=14$  for cations. In  $C_{15}S^+$  and  $C_{16}S^+$ , the cyclic structure is more stable by 0.08 and 0.27 eV, respectively. The bonding in these chains is found to be cumulenic in nature.

Experimentally,  $C_nS$  ( $n=2-9$ ) have been investigated by microwave spectroscopy [92–94]. In addition, IR spectra have been recorded in argon matrices for  $n=1-5$  [95,96]. Information on the excited electronic states in neutral and anionic chains is presented in Table 6, Panel A. However, electronic spectra of the cationic species have not been reported thus far. The  $^1\Sigma^+ \leftarrow X^1\Sigma^+$  transition for  $C_nS$  ( $n=3, 5$ ),  $^3\Sigma^- \leftarrow X^3\Sigma^-$  ( $n=2, 4$

Table 6. S terminated carbon chains.

Species/electronic ground state	$n$	Origin ( $\text{cm}^{-1}$ )	Phase	Reference
<i>Panel A: <math>C_nS</math> species</i>				
$C_nS$	2	$^3\Pi$ 9146	Ne matrix	J. Mol. Spectrosc. <b>222</b> , 15 (2003)
$n$ even $X^3\Sigma^-$		$^3\Sigma^-$ 30,563		
$n$ odd $X^1\Sigma^+$		$^3\Pi$ 9226	Gas	J. Chem. Phys. <b>117</b> , 9327 (2002)
		$^3\Sigma^-$ 30,627		J. Chem. Phys. <b>130</b> , 014302 (2009)
	3	26,460	Ar matrix	Chem. Ber. <b>124</b> , 2617 (1991)
	4	$^3\Sigma^-$ 22,380	Ne matrix	J. Phys. Chem A. <b>105</b> , 4894 (2001)
	5	$^1\Sigma^+$ 35,169	Ne matrix	J. Phys. Chem A <b>107</b> , 8856 (2003)
	6	$^3\Sigma^-$ 17,416	Ne matrix	J. Phys. Chem A. <b>107</b> , 8856 (2003)
$C_nS^-$	2	$^2\Sigma^+$ 10,606	Ne matrix	J. Mol. Spectrosc. <b>222</b> , 15 (2003)
$n$ even $X^2\Pi$		$^2\Pi$ 22,273		
$n$ odd $X^2\Pi$	4	$^2\Pi$ 17,593	Ne matrix	J. Phys. Chem A. <b>105</b> , 4894 (2001)
	6	$^2\Pi$ 16,447	Ne matrix	J. Phys. Chem A. <b>105</b> , 4894 (2001)
<i>Panel B: <math>SC_nS</math> species</i>				
$SC_nS$	2	A 25,510	Ar matrix	J. Am. Chem. Soc. <b>114</b> , 6452 (1992)
$n$ even $X^3\Sigma^-$		B 43,500		
$n$ odd $X^1\Sigma^+$	3	A 430–530 nm	Gas	J. Chem. Phys. <b>56</b> , 5428 (1972)
		B 220–255 nm		
	4	A 19,455	Gas	Chem. Ber. <b>124</b> , 2617 (1991)
		B 36,765		
	5	A 15,100	Gas	Chem. Ber. <b>123</b> , 1753 (1990)
		B 31,250		
$SC_nS^-$	2	$^2\Pi$ 18,423	Gas	Chem. Phys. Lett. <b>410</b> , 172 (2005)
$n$ even $X^2\Pi$				
$n$ odd $X^2\Pi$				
<i>Panel C: <math>HC_nS</math> species</i>				
$HC_nS$	2	$^2\Pi_{3/2}$ 24,197	Gas	J. Chem. Phys. <b>123</b> , 014317 (2005)
$X^2\Pi_{3/2}$ ( $n$ even)	4	$^2\Pi_{3/2}$ 19,981	Gas	Phys. Chem. Chem. Phys. <b>10</b> , 136 (2008)
	6	$^2\Pi_{3/2}$ 16,961	Gas	Int. J. Mass. Spectrom. <b>233</b> , 131 (2004)
	8	$^2\Pi_{3/2}$ 14,838	Gas	Int. J. Mass. Spectrom. <b>233</b> , 131 (2004)
	10	$^2\Pi_{3/2}$ 13,334	Gas	Int. J. Mass. Spectrom. <b>233</b> , 131 (2004)

and 6) and  $A^3\Pi \leftarrow X^3\Sigma^-$  for  $C_2S$  have been observed in 6 K neon matrices. The gas phase LIF spectra of the corresponding band systems in  $C_2S$  have been recorded and rotationally resolved. In the case of the anions,  $B^2\Pi \leftarrow X^2\Pi$  transition for  $n=2, 4$  and  $6$  and the  $A^2\Sigma^+ \leftarrow X^2\Pi$  for  $n=2$  have been observed in neon matrices. In addition,  $C_2S^-$  and  $C_3S^-$  have been investigated via anion photoelectron spectroscopy [77,80]. In the case of  $C_2S$ , two low-lying singlet states,  $a^1\Delta$  and  $b^1\Sigma^+$  have been located at 0.45 and 0.74 eV above the  $X^3\Sigma^-$  ground state. The spin-orbit splitting for the anion ground state,  $X^2\Pi_i$  was around  $195\text{ cm}^{-1}$ . In the case of  $C_3S^-$ , the photoelectron spectrum revealed a strong progression in the  $\nu_4$  bending mode indicating bent and linear equilibrium geometries for the anion and neutral, respectively.

## 6.2. $HC_nS$ , $HC_nS^+$ and $HC_nS^-$

The hydrogenated polycarbon sulphides,  $HC_nS$  are open shell species with  $^2\Pi$  ground states in which the spin-orbit and RT effects play crucial roles towards understanding their electronic structure. The ground state geometries and electronic structure for  $n=2-8$  have been investigated via microwave spectroscopy [97-99]. These studies confirm that the  $HC_nS$  chains have linear structures with the even and odd chains exhibiting  $^2\Pi_i$  ( $\pi^3$ ) and  $^2\Pi_r$  ( $\pi^1$ ) symmetries, respectively, in contrast to the isovalent  $HC_nO$  species which have bent equilibrium geometries. In a recent theoretical study, the spin-orbit coupling in the  $^2\Pi$  ground state for  $HC_nS$  ( $n=1-12$ ) was calculated and determined to be around  $-300\text{ cm}^{-1}$  for the even and  $120\text{ cm}^{-1}$  for the odd numbered chains [100].

In the above series, the even members with  $n=2, 4, 6, 8$  and  $10$  have been investigated by optical spectroscopy in the gas phase (Table 6, Panel C), with  $HC_2S$  being extensively studied both experimentally and theoretically. The LIF spectra of the  $^2\Pi_i \leftarrow X^2\Pi_i$  transition for  $n=2, 4$  and  $6$  have been recorded. In the case of  $HC_2S$ , spectral analysis yielded ground and excited state spin-orbit splitting of  $-254.1$  and  $45.4\text{ cm}^{-1}$ , respectively. In addition, the excited state was found to be quasilinear based on the large RT splitting. Recently, the  $^2\Pi_{3/2} \leftarrow X^2\Pi_{3/2}$  transition of  $HC_2S$  and  $HC_4S$  was recorded in a supersonic slit-jet discharge using degenerate four-wave mixing and two-colour resonant four-wave mixing (TC-RFWM) techniques. Figure 6 depicts the method and assignment procedure for the rotational lines of the TC-RFWM spectrum of  $A^2\Pi_{3/2} \leftarrow X^2\Pi_{3/2}$  transition in  $HC_2S$ . Using this approach, 14 vibrational levels of the  $X^2\Pi_{3/2}$  ground state were rotationally resolved with a high signal-to-noise ratio of  $\sim 10^3$ . The above band system has been recorded for  $HC_6S$  by R2PI and  $HC_8S$  and  $HC_{10}S$  by CRD spectroscopy. A near linear relationship was observed between the absorption wavelength and the chain size, similar to the  $C_nH^-$  chains, suggesting cumulenenic bonding. The band origin of the next higher member,  $HC_{12}S$  has been estimated to be around 820 nm.

Experimental and theoretical data on the corresponding cations and anions is sparse.  $HC_nS$  ions have been generated by laser ablation of a mixture of carbon and sulphur and the products analysed by collision induced dissociation and *ab initio* calculations. The structure of the ions was found to be linear with the hetero atoms located at the ends of the chain. These studies also suggest that the weakest bond is the C-C bond adjacent to the terminal sulphur atom [101].

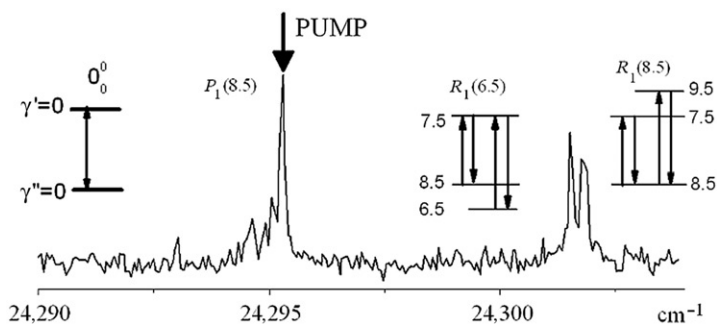


Figure 6. Two-colour resonant four wave mixing (TC-RFWM) spectrum of  $\text{HC}_2\text{S}$  recorded by fixing the pump laser at the  $P_1(8.5)$  transition of the origin band of  $\text{A } ^2\Pi_{3/2} \leftarrow \text{X } ^2\Pi_{3/2}$  band and scanning the pump laser through the  $\text{O}_0^0$  transition. The nearly degenerate TC-RFWM displays three transitions: the degenerate transition at  $P_1(8.5)$  and two  $R_1$  transitions sharing the upper and lower level ( $R_1(6.5)$  and  $R_1(8.5)$ , respectively).

### 6.3. $\text{SC}_n\text{S}$ and $\text{SC}_n\text{S}^-$

The  $\text{SC}_n\text{S}$  chains, like the  $\text{C}_n\text{S}$  species show an even–odd alternation in the ground state symmetry, with the even and odd members represented by  $^3\Sigma^-$  ( $\pi^2$ ) and  $^1\Sigma^+$  ( $\pi^4$ ), respectively. Information on the ground state geometry, vibrational frequencies and the favourable dissociation pathways in clusters up to  $n=29$  have been deduced using the DFT method [96]. The results indicate that these chains have linear ground states and are cumulenic in nature. The dissociation energies of  $\text{SC}_n\text{S}$  with odd  $n$  were higher than the even  $n$ , suggesting that the odd numbered chains are more stable. This is consistent with the complete occupancy of the  $\pi$  orbital in these chains. In addition, the most favoured dissociation channel was the loss of CS species.

Electronic spectra have been recorded for  $n=2-5$  in the vapour phase and argon matrices (Table 6, Panel B). Two-band systems, one in the visible and the other in the UV were observed. In the case of  $\text{C}_3\text{S}_2$ , phosphorescence from the  $^3\Sigma_u^-$  state located around  $18,287\text{ cm}^{-1}$  was observed in argon matrices. Among the anions, LIF spectrum of  $\text{A } ^2\Pi_g \leftarrow \text{X } ^2\Pi_u$  band system in  $\text{C}_2\text{S}_2^-$  has been recently measured. This is the only other anion, besides  $\text{C}_2^-$  [102] and  $\text{C}_4^-$  [103] for which fluorescence measurements are known. From the fluorescence depletion and dispersed emission measurements the C–C and C–S stretching frequencies for the ground and excited states were deduced. The  $\text{A } ^2\Pi$  ( $\nu=0$ ) state was found to have a lifetime of  $\sim 30\text{ ns}$  consistent with the theoretically predicted value of  $20\text{ ns}$  leading to the conclusion that the fluorescence quantum yield is approximately unity for the vibrationless level of the A state.

## 7. Metal terminated carbon chains

Alkaline earth metal acetylides,  $\text{MC}_2\text{H}$ ;  $\text{M}=\text{Mg}, \text{Ca}, \text{Sr}$  are known to have  $^2\Sigma^+$  ground states, where the unpaired electron is located in a sigma orbital which is predominantly metal  $ns$  in character. The electro-positive nature of the alkaline earth metals and the high electron affinity of the  $\text{C}_2\text{H}$  radical (2.9 eV) results in a chemical bond ( $\text{M}-\text{C}_2\text{H}$ ) which can be described as being ionic. The electronic structure and bonding in these systems are

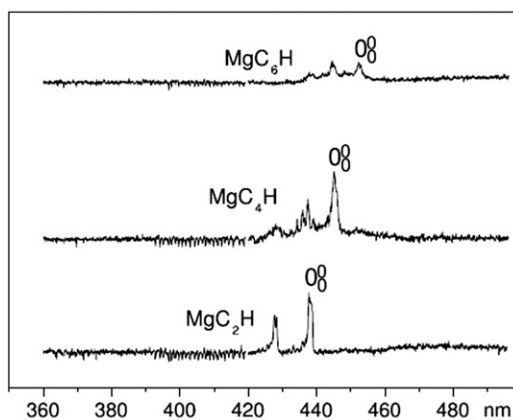


Figure 7. Resonant two-colour, two-photon ionisation spectra of the A  $^2\Pi \leftarrow X ^2\Sigma^+$  electronic transition of  $\text{MgC}_{2n}\text{H}$  ( $n=1-3$ ).

similar to the corresponding well-characterised hydrides, halides and hydroxides. Pure rotational spectra have been recorded for  $\text{MgC}_2\text{H}$  [104],  $\text{CaC}_2\text{H}$  [105] and  $\text{SrC}_2\text{H}$  [106] which confirm a linear  $^2\Sigma^+$  ground state. The lowest electronic states have  $^2\Sigma^+$ ,  $^2\Pi$  and  $^2\Delta$  symmetries with major contributions from the metal  $ns$ ,  $np$  and  $(n-1)d$  orbitals.

The A  $^2\Pi \leftarrow X ^2\Sigma^+$  transition has been observed in the three acetylides and the higher members of the  $\text{MgC}_n\text{H}$  series ( $n=4, 6$ ) as well (Table 7). The  $\pi \leftarrow \sigma$  electron excitation corresponding to the above transition revealed a strong progression in the M–C stretching mode. The stretching frequency was found to be higher for the excited state indicating a decrease in the bond length upon excitation. In addition, the  $f$  values for the A  $^2\Pi \leftarrow X ^2\Sigma^+$  and B  $^2\Pi \leftarrow X ^2\Sigma^+$  transitions calculated using the TD-DFT method for  $\text{MgC}_{2n}\text{H}$  ( $n=1-8$ ) were found to decrease with the chain length [107]. These characteristics have been explained in terms of promotion of an electron from the metal centred  $\sigma$  HOMO which is antibonding in character, to the  $\pi$  LUMO, composed mainly of metal  $np$  orbitals. This results in a stronger M–C bond in the excited state, moreover, as the chain length increases, the  $\sigma$  electron becomes more delocalised resulting in a decrease in the  $f$  value. The REMPI spectrum of the A  $^2\Pi \leftarrow X ^2\Sigma^+$  electronic transition of  $\text{MgC}_{2n}\text{H}$  ( $n=1-3$ ) is shown in Figure 7. The B  $^2\Sigma^+ \leftarrow X ^2\Sigma^+$  ( $\text{CaC}_2\text{H}$ ) and the orbitally forbidden C  $^2\Delta \leftarrow X ^2\Sigma^+$  ( $\text{CaC}_2\text{H}$  and  $\text{SrC}_2\text{H}$ ) transitions have been observed by CRDS and LIF. In the case of  $\text{AlC}_2\text{H}$ , DFT calculations predicted two structures; the linear and cyclic isomers, with the former being stable by  $41 \text{ kJ mol}^{-1}$ . Recently, aluminium monoacetylide was investigated using a REMPI technique. The rotationally resolved spectrum corresponding to the origin band is displayed in Figure 8. Based on the analysis, the spectrum was assigned to the A  $^1\Pi \leftarrow X ^1\Sigma^+$  transition of linear  $\text{AlC}_2\text{H}$ .

The other metal terminated carbon chains with known optical transitions include  $\text{AlC}_2$  and  $\text{SiC}_2$  with the latter having been detected in IRC + 10216 [108,109]. By analogy to  $\text{C}_3$ ,  $\text{SiC}_2$  was thought to possess a linear structure. However, rotationally resolved R2PI studies confirmed a T-shaped cyclic structure ( $\text{C}_{2v}$ ); a three-membered ring with  $\text{C}\equiv\text{C}$  triple bond corresponding to a X  $^1\text{A}_1$  ground state. The linear isomer ( $\text{C}_{\infty v}$ ) has been predicted to lie around  $25 \text{ kJ mol}^{-1}$  higher in energy [110]. Among the larger members,

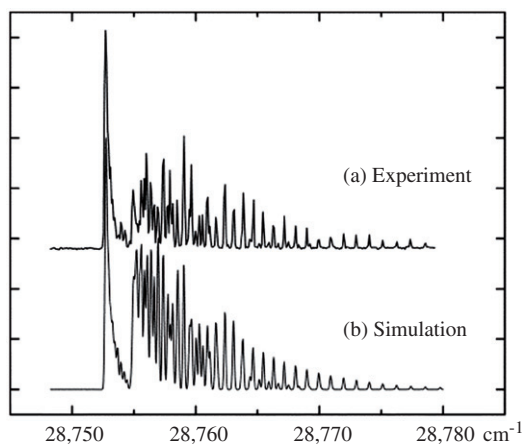


Figure 8. (a) Experimental and (b) simulated spectra corresponding to the origin band of the  $A^1\Pi \leftarrow X^1\Sigma^+$  electronic transition for AICCH measured using the resonant two-colour two-photon ionisation technique.

cyclic  $\text{SiC}_3$  [111] and linear  $\text{SiC}_4$  [112] have been detected in the circumstellar media. Pure rotational spectra of linear  $\text{SiC}_n$  ( $n = 3$  and  $5-8$ ) chains have been recorded which confirm an even/odd alternation in the ground state symmetry;  $^1\Sigma^+$  for the even  $n$  chains and  $^3\Sigma^-$  for the odd  $n$  systems [113]. The corresponding anions  $\text{SiC}_n^-$  ( $n = 3-7$ ) have been investigated by photoelectron spectroscopy and are shown to possess a linear  $C_{\infty v}$  ground state geometry [114,115]. Like silicon dicarbide,  $\text{AlC}_2$  exhibits a cyclic structure ( $X^2A_1$ ) which is more stable than the linear ( $X^2\Sigma^+$ ) isomer by  $33 \text{ kJ mol}^{-1}$  [116] ( $46 \text{ kJ mol}^{-1}$ ) [117]. In the case of  $\text{AlC}_2$ , bonding has been explained in terms of electron donation from the more electropositive Al to the electronegative  $\text{C}_2$  via  $\sigma$  and  $\pi$  coordination. The first excited  $A^1B_2$  state has been experimentally observed for  $\text{SiC}_2$  and higher electronic states with  $^1A_1$  and  $^1B_1$  symmetries have been predicted. The lowest excited  $^2A_1$  in  $\text{AlC}_2$  was detected around 1260 nm by photoelectron studies. Recently, the  $C^2B_2 \leftarrow X^1A_1$  and  $D^2B_1 \leftarrow X^1A_1$  were rotationally analysed by LIF and REMPI.

## 8. Outlook

Over the years, researchers have employed a variety of experimental techniques to record electronic spectra of carbon chains and ions. This has created a wealth of data which continues to grow. Initial measurements carried out on mass selected species condensed in argon or neon matrices have been the guiding tools for further gas phase measurements. Spectra recorded in the gas phase using supersonic expansion or by buffer gas cooling of trapped ions are associated with internally cold molecules and are also free of matrix perturbations. This is crucial for direct comparisons with astrophysical observations. The above methodology has been successfully implemented on bare carbon chains, their hydrogen and nitrogen derivatives leading to a general consensus that C-chains with less than 10 carbon atoms cannot be carriers of DIB absorptions that lie in the 400–900 nm range. Instead, closed shell species ( $^1\Sigma^+$ ) such as the longer odd carbon chains

Table 7. Metal terminated carbon chains:  $MC_nH/MC_n$  species.

Species/electronic ground state	$n$	Origin ( $\text{cm}^{-1}$ )	Phase	Reference
$MgC_nH$	2	$A^2\Pi$ 22,809	Gas	J. Mol. Spectrosc. <b>230</b> , 54 (2005)
$X^2\Sigma^+$	4	$A^2\Pi$ 22,450	Gas	ApJ. <b>677</b> , 348 (2008)
	6	$A^2\Pi$ 22,093	Gas	ApJ. <b>677</b> , 348 (2008)
$CaC_nH$	2	$A^2\Pi_{1/2}$ 15,487	Gas	Chem. Phys. Lett. <b>136</b> , 97 (1987)
$X^2\Sigma^+$		$A^2\Pi_{3/2}$ 15,560		
		$B^2\Sigma^+$ 16,567	Gas	Chem. Phys. Lett. <b>433</b> , 264 (2007)
		$C^2\Delta$ 22,157	Gas	Chem. Phys. Lett. <b>288</b> , 785 (1998)
$SrC_nH$	2	$A^2\Pi_{1/2}$ 14,176	Gas	Chem. Phys. Lett. <b>136</b> , 97 (1987)
$X^2\Sigma^+$		$A^2\Pi_{3/2}$ 15,560		
		$B^2\Delta_{3/2}$ 21,077	Gas	J. Mol. Spectrosc. <b>206</b> , 198 (2001)
		$B^2\Delta_{5/2}$ 21,183		
$AlC_nH$	2	$A^1\Pi$ 28,750	Gas	Phys. Chem. Chem. Phys. <b>9</b> , 3897 (2007)
$X^1\Sigma^+$				
$AlC_n$	2	$A^2A_1$ 7910	Gas	J. Am. Chem. Soc. <b>121</b> , 10193 (1999)
$X^2A_1$		$C^2B_2$ 22,098		J. Chem. Phys. <b>131</b> , 064305 (2009)
		$D^2B_1$ 26,073		
$SiC_n$	2	$A^1B_2$ 20,080	Gas	J. Chem. Phys. <b>80</b> , 3556 (1984)
$X^1A_1$				

$C_{17}, C_{19}, \dots, C_{2n+1}H^+, C_{2n}N^+, C_{2n+1}O$  and related molecules which have strong  $^1\Sigma^+ \leftarrow X^1\Sigma^+$  transitions in the DIB range remain as viable candidates. Gas phase spectroscopy of sulphur bearing carbon chains isovalent with the  $C_nO$  systems and metal terminated (Si, Fe, Mg and Al) carbon chains would be a worthwhile endeavour.

### Acknowledgement

This research is supported by the Swiss National Science Foundation (200020 – 124349/1).

### References

- [1] A. Belloche, R. T. Garrod, H. S. P. Mueller, K. M. Menten, C. Comito, and P. Schilke, *Astron. Astrophys.* **499**, 215 (2009).
- [2] L. M. Ziurys, *Proc. Natl. Acad. Sci. USA.* **103**, 12274 (2006).
- [3] D. A. Williams, *Faraday Discuss.* **109**, 1 (1998).
- [4] E. Herbst, *Annu. Rev. Phys. Chem.* **46**, 27 (1995).
- [5] R. I. Kaiser, C. Ochsenfeld, D. Stranges, M. Head-Gordon, and Y. T. Lee, *Faraday Discuss.* **109**, 183 (1998).
- [6] G. H. Herbig, *Astrophys. J.* **196**, 129 (1975).
- [7] A. G. G. M. Tielens and T. P. Snow, editors, *The Diffuse Interstellar Bands* (Kluwer, Dordrecht, 1995).
- [8] J. P. Maier, *Chem. Soc. Rev.* **26**, 21 (1997).
- [9] E. B. Jochnowitz and J. P. Maier, *Annu. Rev. Phys. Chem.* **59**, 519 (2008).



- [10] J. Cernicharo, M. Guelin, M. Agundez, K. Kawaguchi, M. McCarthy, and P. Thaddeus, *Astron. Astrophys.* **467**, L37 (2007).
- [11] M. C. McCarthy, C. A. Gottlieb, H. Gupta, and P. Thaddeus, *Astrophys. J.* **652**, L141 (2006).
- [12] S. Brunken, H. Gupta, C. A. Gottlieb, M. C. McCarthy, and P. Thaddeus, *Astrophys. J.* **664**, L43 (2007).
- [13] K.-H. Homann, *Angew. Chem., Int. Ed.* **37**, 2435 (1998).
- [14] J. G. Kushmerick, J. Naciri, J. C. Yang, and R. Shashidhar, *Nano Lett.* **3**, 897 (2003).
- [15] N. Robertson and C. A. McGowan, *Chem. Soc. Rev.* **32**, 96 (2003).
- [16] P. Thaddeus and M. C. McCarthy, *Spectrochim. Acta* **57A**, 757 (2001).
- [17] K. H. Hinkle, *Lect. Notes Phys.* **428**, 98 (1994).
- [18] P. Botschwina, *Phys. Chem. Chem. Phys.* **5**, 3337 (2003).
- [19] C. Liang and H. F. Schaefer, III, *J. Chem. Phys.* **93**, 8844 (1990); V. Parasuk and J. Almlof, *J. Chem. Phys.* **94**, 8172 (1991).
- [20] J. M. L. Martin and P. R. Taylor, *J. Phys. Chem.* **100**, 6047 (1996).
- [21] T. F. Giesen, A. Van Orden, H. J. Hwang, R. S. Fellers, R. A. Provencal, and R. J. Saykally, *Science* **265**, 756 (1994).
- [22] J. R. Heath and R. J. Saykally, in *On Clusters and Clustering*, edited by P.J. Reynolds (Elsevier, Amsterdam, 1993).
- [23] J. D. Presilla-Marquez, J. A. Sheehy, J. D. Mills, P. G. Carrick, and C. W. Larson, *Chem. Phys. Lett.* **274**, 439 (1997).
- [24] S. L. Wang, C. M. L. Rittby, and W. R. M. Graham, *J. Chem. Phys.* **107**, 7025 (1997).
- [25] S. L. Wang, C. M. L. Rittby, and W. R. M. Graham, *J. Chem. Phys.* **107**, 6032 (1997).
- [26] M. Grutter, M. Wyss, E. Riaplov, J. P. Maier, S. D. Peyerimhoff, and M. Hanrath, *J. Chem. Phys.* **111**, 7397 (1999).
- [27] A. E. Boguslavskiy, H. Ding, and J. P. Maier, *J. Chem. Phys.* **123**, 34305 (2005).
- [28] A. E. Boguslavskiy and J. P. Maier, *Phys. Chem. Chem. Phys.* **9**, 127 (2007).
- [29] M. T. Bowers, P. R. Kemper, G. von Helden, and P. A. M. van Koppen, *Science* **260**, 1446 (1993).
- [30] N. G. Gotts, G. von Helden, and M. T. Bowers, *Int. J. Mass Spectrom. Ion Processes* **149/150**, 217 (1995).
- [31] D. C. Parent and S. L. Anderson, *Chem. Rev.* **92**, 1541 (1992).
- [32] T.P. Snow and B. J. McCall, *Annu. Rev. Astron. Astrophys.* **44**, 367 (2006).
- [33] L. M. Haffner and D. M. Meyer, *Astrophys. J.* **453**, 450 (1995).
- [34] P. F. Bernath, K. H. Hinkle, and J. J. Keady, *Science* **244**, 562 (1989).
- [35] D. W. Arnold, S. E. Bradforth, T. N. Kitsopoulos, and D. M. Neumark, *J. Chem. Phys.* **95**, 8753 (1991).
- [36] M. Guelin, J. Cernicharo, M. J. Travers, M. C. McCarthy, C. A. Gottlieb, P. Thaddeus, M. Ohishi, S. Saito, and S. Yamamoto, *Astron. Astrophys.* **317**, L1 (1997).
- [37] C. A. Gottlieb, M. C. McCarthy, M. J. Travers, J. U. Grabow, and P. Thaddeus, *J. Chem. Phys.* **109**, 5433 (1998).
- [38] T. D. Crawford, J. F. Stanton, J. C. Saeh, and H. F. Schaefer III, *J. Am. Chem. Soc.* **121**, 1902 (1999).
- [39] J. Takahashi, *Publ. Astron. Soc. Jpn.* **52**, 401 (2000).
- [40] C. Ochsenfeld, R. I. Kaiser, Y. T. Lee, A. G. Suits, and M. Head-Gordon, *J. Chem. Phys.* **106**, 4141 (1997).
- [41] T. R. Taylor, C. Xu, and D. M. Neumark, *J. Chem. Phys.* **108**, 10018 (1998).
- [42] C. Desfrancois, H. Abdoul-Carime, and J.-P. Schermann, *Int. J. Mod. Phys. B* **10**, 1339 (1996).
- [43] F. Pauzat, Y. Ellinger, and A. D. McLean, *Astrophys. J.* **369**, L13 (1991).
- [44] S. Graf, J. Geiss, and S. Leutwyler, *J. Chem. Phys.* **114**, 4542 (2001).
- [45] D. E. Woon, *Chem. Phys. Lett.* **244**, 45 (1995).
- [46] S. J. Kim, J. Caldwell, A. R. Rivolo, R. Wagener, and G. S. Orton, *Icarus* **64**, 233 (1985).

- [47] W. D. Langer, T. Velusamy, T. B. H. Kuiper, R. Peng, M. C. McCarthy, M. J. Travers, A. Kovacs, C. A. Gottlieb, and P. Thaddeus, *Astrophys. J.* **480**, L63 (1997).
- [48] V. G. Kunde, A. C. Aikin, R. A. Hanel, D. E. Jennings, W. C. Maguire, and R. E. Samuelson, *Nature* **292**, 686 (1981).
- [49] A. Scemama, P. Chaquin, M. -C. Gazeau, and Y. Benilan, *J. Phys. Chem. A* **106**, 3828 (2002).
- [50] A. Mavrandonakis, M. Muehlhaeuser, G. E. Froudakis, and S. D. Peyerimhoff, *Phys. Chem. Chem. Phys.* **4**, 3318 (2002).
- [51] G. Mpourmpakis, M. Muhlhauser, G. E. Froudakis, and S. D. Peyerimhoff, *Chem. Phys. Lett.* **356**, 398 (2002).
- [52] H. Ding, T. W. Schmidt, T. Pino, A. E. Boguslavskiy, F. Guthe, and J. P. Maier, *J. Chem. Phys.* **119**, 814 (2003).
- [53] M. Guelin and P. Thaddeus, *Astrophys. J.* **212**, L81 (1977).
- [54] M. Guelin, N. Neininger, and J. Cernicharo, *Astron. Astrophys.* **335**, L1 (1998).
- [55] P. Botschwina, M. Horn, K. Markey, and R. Oswald, *Mol. Phys.* **92**, 381 (1997).
- [56] J. J. Belbruno, Z.-C. Tang, R. Smith, and S. Hobday, *Mol. Phys.* **99**, 957 (2001).
- [57] P. Thaddeus, C. A. Gottlieb, H. Gupta, S. Brunken, M. C. McCarthy, M. Agundez, M. Guelin, and J. Cernicharo, *Astrophys. J.* **677**, 1132 (2008).
- [58] J. Cernicharo, M. Guelin, M. Agundez, M. C. McCarthy, and P. Thaddeus, *Astrophys. J.* **688**, L83 (2008).
- [59] E. Garand, T. I. Yacovitch, and D. M. Neumark, *J. Chem. Phys.* **130**, 064304/1 (2009).
- [60] M. C. McCarthy, J. U. Grabow, M. J. Travers, W. Chen, C. A. Gottlieb, and P. Thaddeus, *Astrophys. J.* **494**, L231 (1998).
- [61] M. B. Bell, P. A. Feldman, M. J. Travers, M. C. McCarthy, C. A. Gottlieb, and P. Thaddeus, *Astrophys. J.* **483**, L61 (1997).
- [62] S. Ikuta, T. Tsuboi, and K. Aoki, *J. Mol. Struct.: THEOCHEM* **528**, 297 (2000).
- [63] K. Aoki, S. Ikuta, and A. Murakami, *Chem. Phys. Lett.* **209**, 211 (1993).
- [64] M. C. McCarthy, J. U. Grabow, M. J. Travers, W. Chen, C. A. Gottlieb, and P. Thaddeus, *Astrophys. J.* **513**, 305 (1999).
- [65] J. Tang, Y. Sumiyoshi, and Y. Endo, *Chem. Phys. Lett.* **315**, 69 (1999).
- [66] Z.-Y. Jiang, X.-H. Xu, H.-S. Wu, F.-Q. Zhang, and Z.-H. Jin, *Int. J. Quantum Chem.* **97**, 876 (2004).
- [67] B. E. Turner and J. Bally, *Astrophys. J.* **321**, L75 (1987).
- [68] L. M. Ziurys, *Astrophys. J.* **321**, L81 (1987).
- [69] M. Guelin, J. Cernicharo, G. Paubert, and B. E. Turner, *Astron. Astrophys.* **230**, L9 (1990).
- [70] D. T. Halpern, D. J. Clouthier, and L. M. Ziurys, *Astrophys. J.* **677**, L101 (2008).
- [71] G. Pascoli and H. Lavendy, *J. Phys. Chem. A* **103**, 3518 (1999).
- [72] M. Ohishi, H. Suzuki, S. I. Ishikawa, C. Yamada, H. Kanamori, W. M. Irvine, R. D. Brown, P. D. Godfrey, and N. Kaifu, *Astrophys. J.* **380**, L39 (1991).
- [73] H. E. Matthews, W. M. Irvine, P. Friberg, R. D. Brown, and P. D. Godfrey, *Nature* **310**, 125 (1984).
- [74] R. D. Brown, P. D. Godfrey, P. S. Elmes, M. Rodler, and L. M. Tack, *J. Am. Chem. Soc.* **107**, 4112 (1985).
- [75] Y. Ohshima, Y. Endo, and T. Ogata, *J. Chem. Phys.* **102**, 1493 (1995).
- [76] T. Ogata, Y. Ohshima, and Y. Endo, *J. Am. Chem. Soc.* **117**, 3593 (1995).
- [77] E. Garand, T. I. Yacovitch, and D. M. Neumark, *J. Chem. Phys.* **129**, 074312/1 (2008).
- [78] J. M. Oakes and G. B. Ellison, *Tetrahedron* **42**, 6263 (1986).
- [79] V. Zengin, B. J. Persson, K. M. Strong, and R. E. Continetti, *J. Chem. Phys.* **105**, 9740 (1996).
- [80] E. Garand, T. I. Yacovitch, and D. M. Neumark, *J. Chem. Phys.* **131**, 054312/1 (2009).
- [81] J. V. Michael and A. F. Wagner, *J. Phys. Chem.* **94**, 2453 (1990).
- [82] B. E. Turner and T. J. Sears, *Astrophys. J.* **340**, 900 (1989).
- [83] Y. Endo and E. Hirota, *J. Chem. Phys.* **86**, 4319 (1987).

- [84] A. L. Cooksy, J. K. G. Watson, C. A. Gottlieb, and P. Thaddeus, *J. Chem. Phys.* **101**, 178 (1994).
- [85] S. Mohamed, M. C. McCarthy, A. L. Cooksy, C. Hinton, and P. Thaddeus, *J. Chem. Phys.* **123**, 234301 (2005).
- [86] D. Strel'nikov, R. Reusch, and W. Kratschmer, *J. Mol. Spectrosc.* **243**, 189 (2007).
- [87] Y. Hirahara, H. Suzuki, S. Yamamoto, K. Kawaguchi, N. Kaifu, M. Ohishi, S. Takano, S. Ishikawa, and A. Masuda, *Astrophys. J.* **394**, 539 (1992).
- [88] M. B. Bell, L. W. Avery, and P. A. Feldman, *Astrophys. J.* **417**, L37 (1993).
- [89] J. Cernicharo, M. Guélin, H. Hein, and C. Kahane, *Astron. Astrophys.* **181**, L9 (1987).
- [90] G. Pascoli and H. Lavendy, *Int. J. Mass Spectrom.* **181**, 11 (1998).
- [91] Z. Tang and J. J. BelBruno, *Int. J. Mass Spectrom.* **208**, 7 (2001).
- [92] S. Yamamoto, S. Saito, K. Kawaguchi, N. Kaifu, H. Suzuki, and M. Ohishi, *Astrophys. J.* **317**, L119 (1987).
- [93] S. Saito, K. Kawaguchi, S. Yamamoto, M. Ohishi, H. Suzuki, and N. Kaifu, *Astrophys. J.* **317**, L115 (1987).
- [94] V. D. Gordon, M. C. McCarthy, A. J. Apponi, and P. Thaddeus, *Astrophys. J. Suppl. Ser.* **134**, 311 (2001).
- [95] H. Wang, J. Szczepanski, A. Cooke, P. Brucat, and M. Vala, *Int. J. Quantum Chem.* **102**, 806 (2005).
- [96] H. Wang, J. Szczepanski, P. Brucat, and M. Vala, *Int. J. Quantum Chem.* **102**, 795 (2005).
- [97] E. Kim, H. Habara, and S. Yamamoto, *J. Mol. Spectrosc.* **212**, 83 (2002).
- [98] Y. Hirahara, Y. Ohshima, and Y. Endo, *J. Chem. Phys.* **101**, 7342 (1994).
- [99] V. D. Gordon, M. C. McCarthy, A. J. Apponi, and P. Thaddeus, *Astrophys. J. Suppl. Ser.* **138**, 297 (2002).
- [100] A. Mitrushchenkov, R. Linguerrri, P. Rosmus, and J. P. Maier, *Mol. Phys.* **107**, 1549 (2009).
- [101] Z.-Y. Liu, Z.-C. Tang, R.-B. Huang, Q. Zhang, and L.-S. Zheng, *J. Phys. Chem. A* **101**, 4019 (1997).
- [102] S. Leutwyler, J. P. Maier, and L. Misev, *Chem. Phys. Lett.* **209**, 211 (1993).
- [103] M. Schaefer, M. Grutter, J. Fulara, D. Forney, P. Ferivogel, and J. P. Maier, *Chem. Phys. Lett.* **209**, 211 (1993).
- [104] M. A. Anderson and L. M. Ziurys, *Astrophys. J.* **439**, L25 (1995).
- [105] M. A. Anderson and L. M. Ziurys, *Astrophys. J.* **444**, L57 (1995).
- [106] B. P. Nuccio, A. J. Apponi, and L. M. Ziurys, *Chem. Phys. Lett.* **247**, 283 (1995).
- [107] H. Ding, C. Apetrei, L. Chacaga, and J. P. Maier, *Astrophys. J.* **77**, 348 (2008).
- [108] P. Thaddeus, S. E. Cummins, and R. A. Linke, *Astrophys. J.* **290**, L29 (1985).
- [109] L. E. Snyder, C. Henkel, J. M. Hollis, and F. J. Lovas, *Astrophys. J.* **290**, L29 (1985).
- [110] I. M.B. Nielsen, W. D. Allen, A. G. Csaszar, and H. F. Schaefer III, *J. Chem. Phys.* **107**, 1195 (1997).
- [111] A. J. Apponi, M. C. McCarthy, C. A. Gottlieb, and P. Thaddeus, *Astrophys. J.* **516**, L103 (1999).
- [112] M. Ohishi, N. Kaifu, K. Kawaguchi, A. Murakami, S. Saito, S. Yamamoto, S. Ishikawa, Y. Fujita, Y. Shiratori, and W. M. Irvine, *Astrophys. J.* **345**, L83 (1989).
- [113] M. C. McCarthy, A. J. Apponi, C. A. Gottlieb, and P. Thaddeus, *Astrophys. J.* **538**, 766 (2000).
- [114] A. Nakajima, T. Taguwa, K. Nakao, M. Gomei, R. Kishi, S. Iwata, and K. Kaya, *J. Chem. Phys.* **103**, 2050 (1995).
- [115] G. E. Davico, R. L. Schwartz, and W. C. Lineberger, *J. Chem. Phys.* **115**, 1789 (2001).
- [116] G. V. Chertihin, L. Andrews, and P. R. Taylor, *J. Am. Chem. Soc.* **116**, 3513 (1994).
- [117] J. R. Flores and A. Largo, *Chem. Phys.* **140**, 19 (1990).